

# ***Theory encounters experiments of Mössbauer spectroscopy***

## ***Tutorial lecture***

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***International Conference on the Applications of the Mössbauer  
Effect***

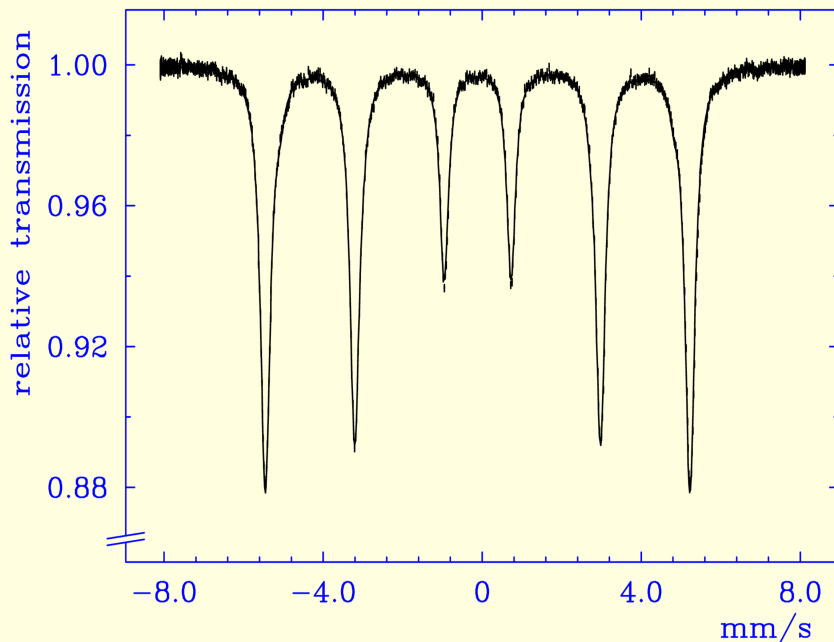
***Saint Petersburg, Russia, 3-8 September 2017***

# **Spectrum**

## ***57Co/Rh source, $\alpha$ -iron (25 $\mu$ m) absorber***

### **Theory**

#### ***H, Is, (texture)***



**$\chi^2$  value**  
***(close to expectation)?***

***Line width/shape?***

***area/thickness =>25 $\mu$ m ?***

***7 points concerning the***  
***Spectrometer/Adaption of theory to experiment***

# Outline

- ***Reduced  $\chi^2$***
- ***Baseline of a transmission spectrum***
- ***The raw data problem***
- ***Choice of the drive frequency***
- ***Theoretical value at channel  $i$***
- ***Dead time***
- ***Background***
  
- ***Conclusion: some good practices***

## **Reduced $\chi^2$**

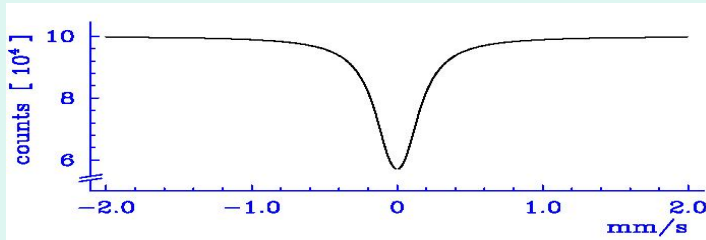
*Counts: Poisson distribution:*  $\chi^2 = 1.0 \pm \sqrt{\frac{2}{N_{ch}}}$ ,  $N_{ch} = 2048 \Rightarrow \chi^2 = 1 \pm 0.031$

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$$N_{ch} = 2048 \Rightarrow \chi^2 = 1 \pm 0.031$$

$$t_{eff} = 10$$



*Beer-Lambert*

law

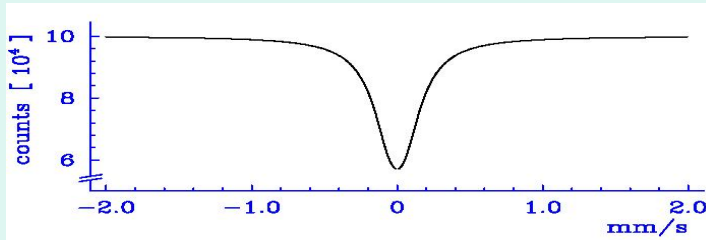
$$Sp(\nu) \propto \int_{-\infty}^{\infty} L_S(\tau - \nu) e^{(-t_{eff} L_A(\tau))} d\tau$$

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**Fit: thin absorber approximation**

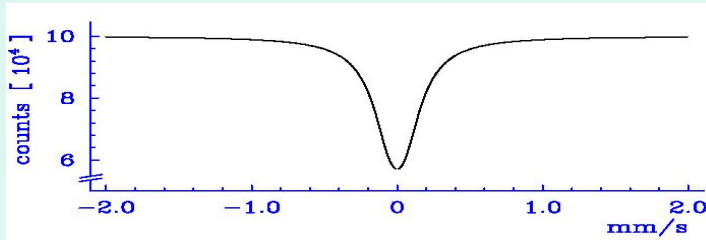
$$Sp(\nu) \propto 1 - t_{eff} L(\nu, \Gamma = \Gamma_S + \Gamma_A)$$

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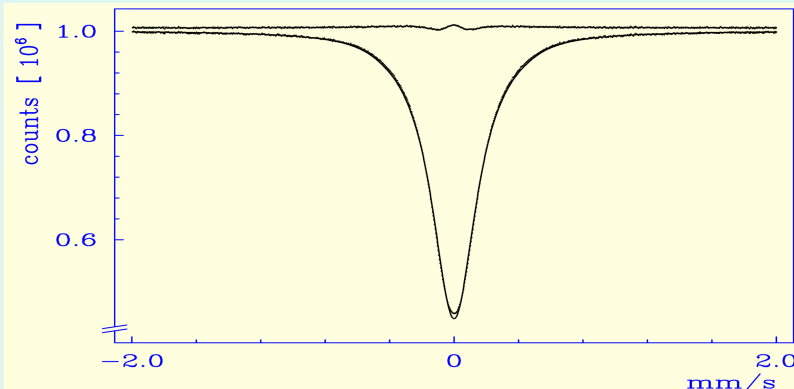
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## Fit: thin absorber approximation

$$Sp(\nu) \propto 1 - t_{eff} L(\nu, \Gamma = \Gamma_S + \Gamma_A)$$

### **synthetic data by Monte Carlo simulations**

(AS70 algorithm of Odeh and Evans, 1974)



*Lorentz-curve*  $\chi^2 = 5.47, \Gamma = 2.49\Gamma_N$

*Voigt profile*  $\chi^2 = 1.42, \Gamma = 2.15\Gamma_N, \sigma_{Gauss} = 1.17\Gamma_N$

# Reduced $\chi^2$

## Note:

### Fit: thin absorber approximation

area is obtained to a good approximation  
by the  $\chi^2$ -fit procedure

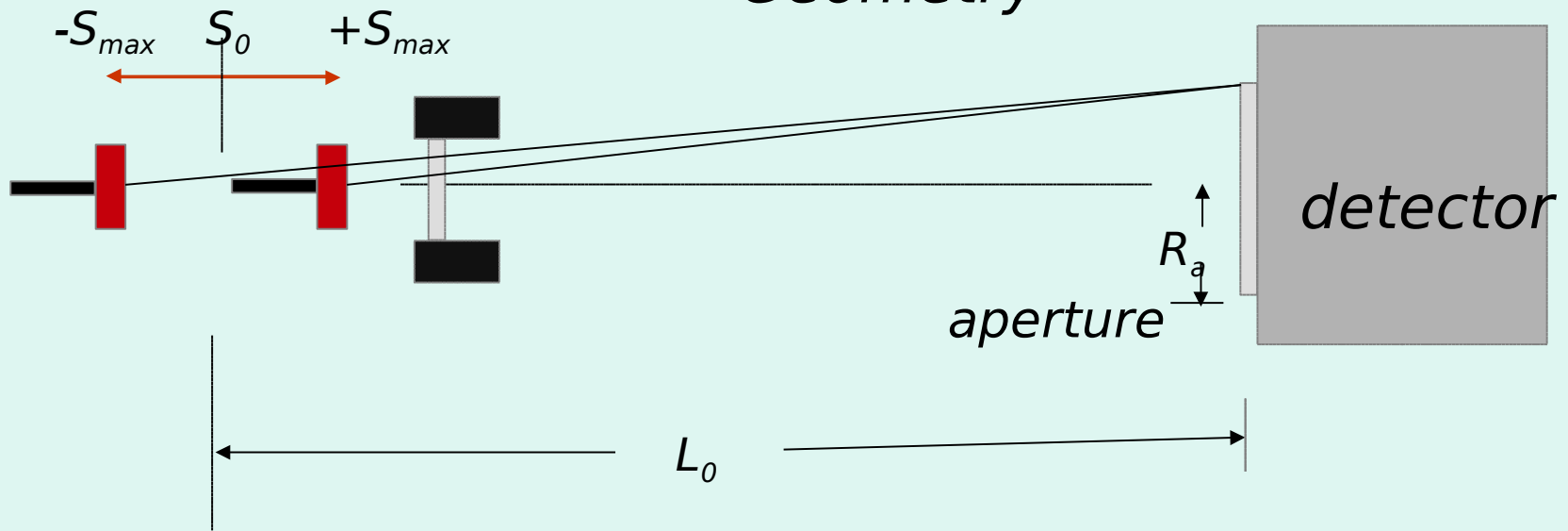
### Fit: finite thickness

Use of Beer-Lambert law and convolution integral:  
thickness instead of area  
 $\chi^2$ -value gets a meaning (validity of the theory)



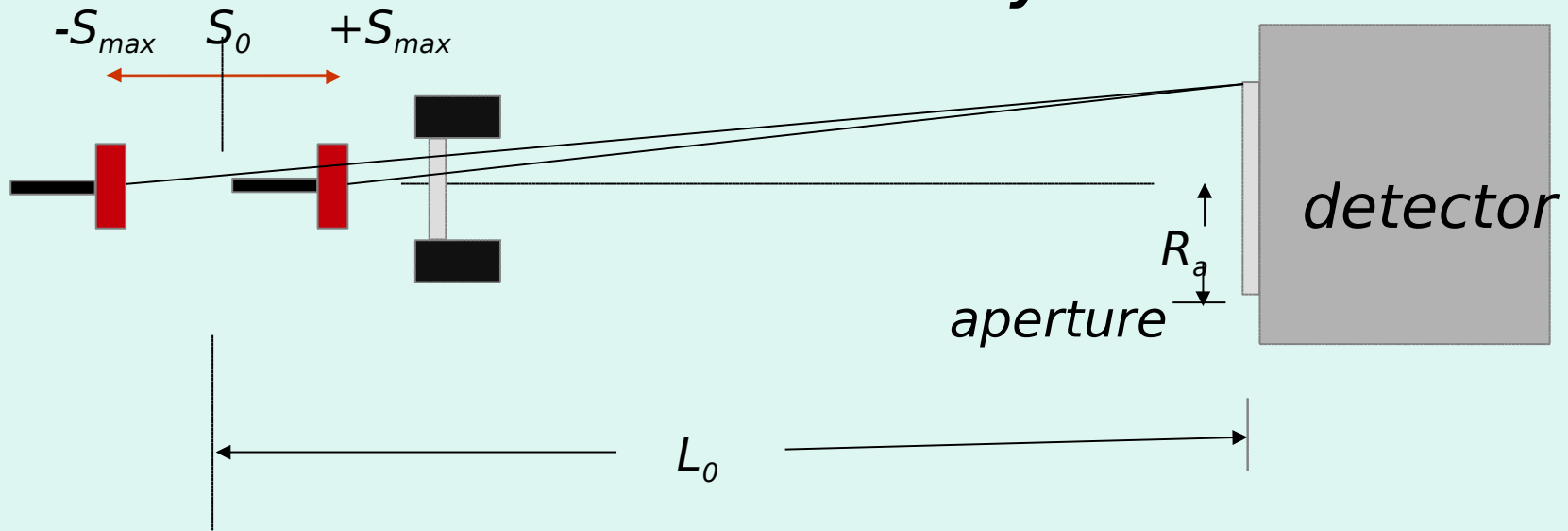
# Baseline of a transmission spectrum

Geometry

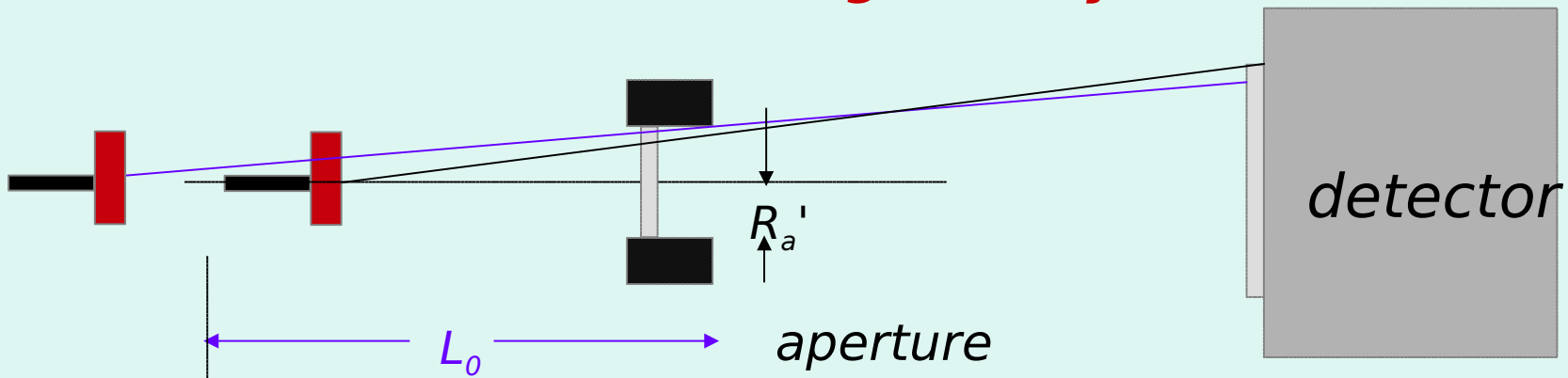


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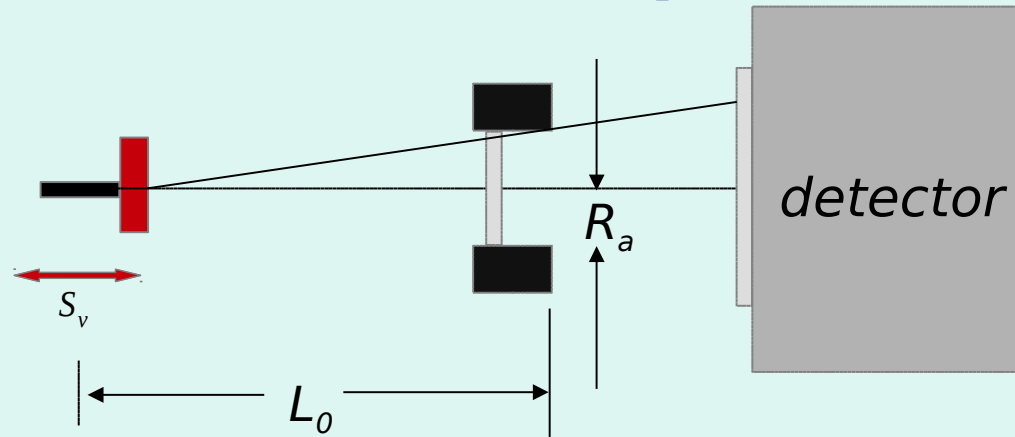
## Geometry



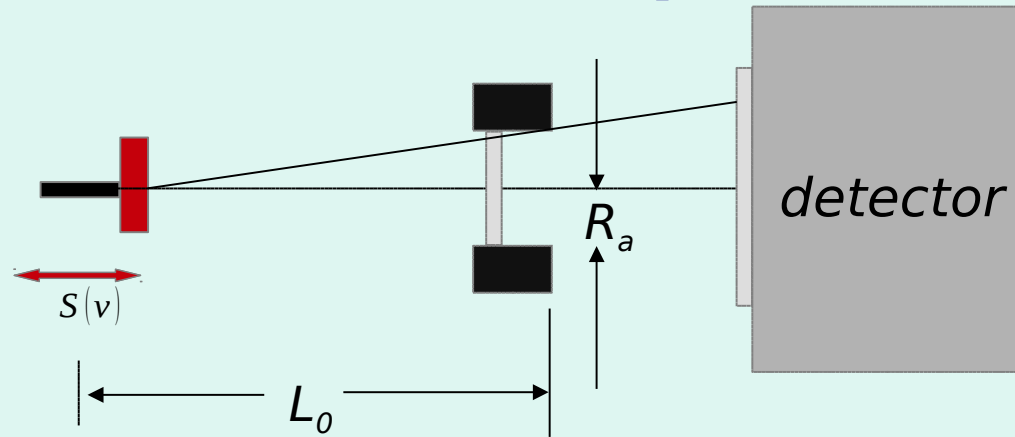
## Undefined geometry



# ***Baseline of a transmission spectrum***

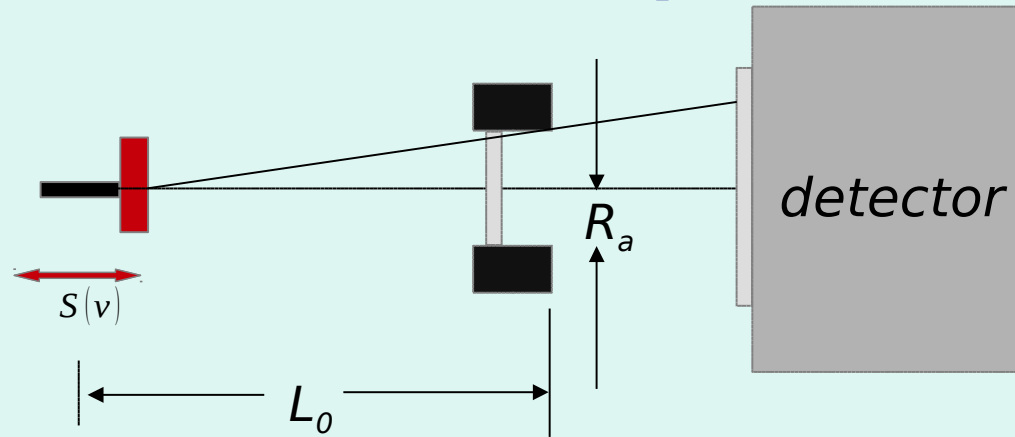


# Baseline of a transmission spectrum



Solid angle:  $\Omega_0 = 2\pi \left( 1 - \frac{L_0}{\sqrt{L_0^2 + R_a^2}} \right)$ ,  $\Omega(\nu) = 2\pi \left( 1 - \frac{L_0 + S(\nu)}{\sqrt{(L_0 + S(\nu))^2 + R_a^2}} \right)$

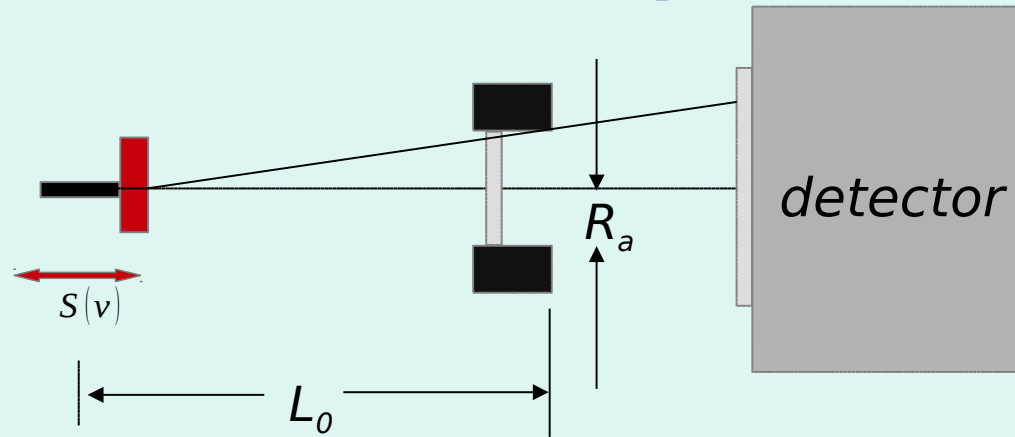
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Count rate  $(\nu)$ :  $\frac{\Omega(\nu)}{\Omega_0} = \frac{1}{(1+\delta)^2} \left( 1 + \frac{3}{4} \varepsilon^2 \left( 1 - \frac{1}{(1+\delta)^2} \right) + \dots \right) \approx \frac{1}{(1+\delta)^2}$ ,  $\varepsilon = \frac{R_a}{L_0}$ ,  $\delta(\nu) = \frac{S(\nu)}{L_0}$

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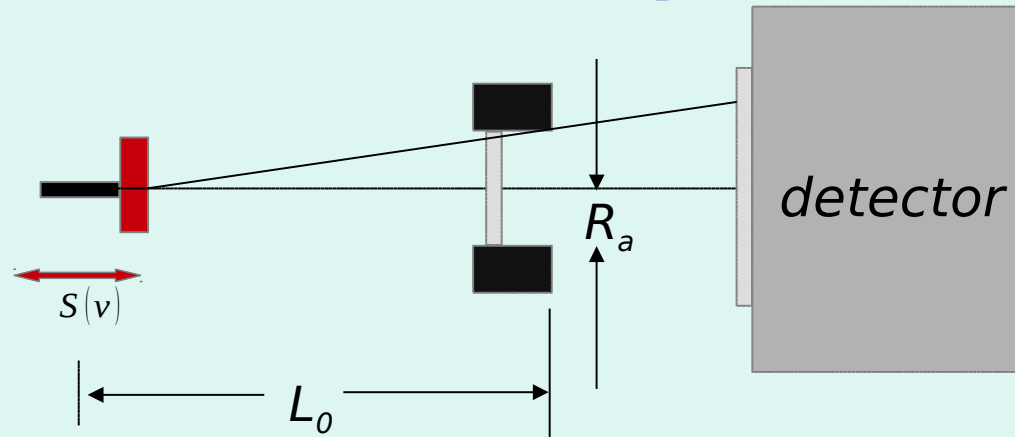


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$S(t) = S_{max} \sin(\omega t)$ ,  $\nu(t) = S_{max} \omega \cos(\omega t)$   $\rightarrow S_{max} \omega = \nu_{max}$ ,  $\frac{S(v)}{L_0} = \pm \frac{S_{max}}{L_0} \sqrt{1 - \left(\frac{\nu}{\nu_{max}}\right)^2}$

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**Sinusoidal:**  $S_{max} = \frac{\nu_{max}}{2\pi f}$

**Triangular:**  $S_{max} = \frac{\nu_{max}}{8f}$

# ***Baseline of a transmission spectrum***

## **Note:**

*The baseline is determined by three parameters:*

$$\text{counts} (v = \infty), \quad \text{geo} = \frac{S_{\max}}{L_0}, \quad \text{channel}_{v=0}$$

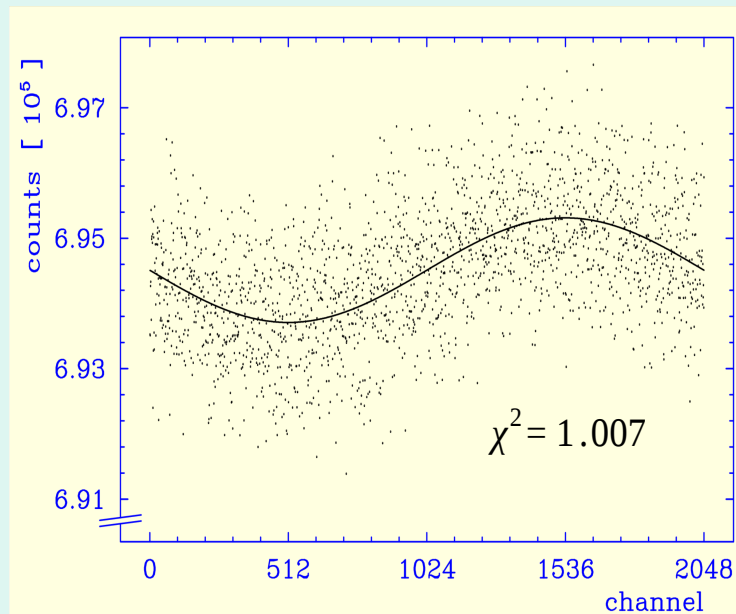


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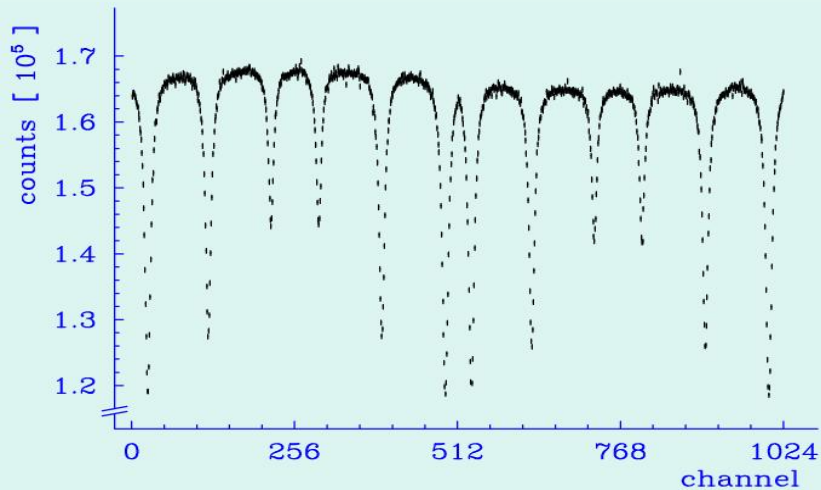
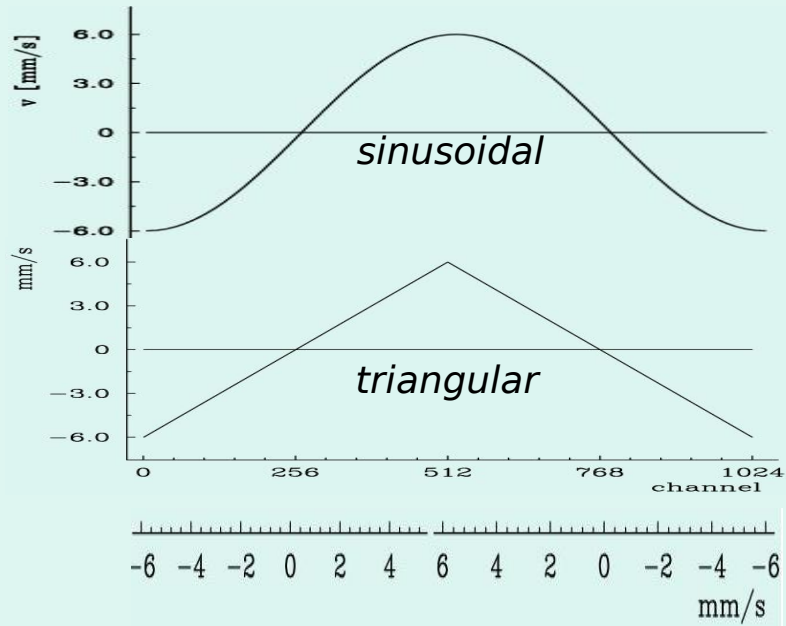
$$\text{geo} = 5.77 \cdot 10^{-4} \quad (\text{fit!})$$

$$\text{geo} = \frac{v_{\max}}{2\pi L_0 f} = \frac{7.24 \text{ mm/s}}{2\pi \cdot 120 \text{ mm} \cdot 17 \text{ Hz}}$$

$$= 5.65 \cdot 10^{-4}$$

(120 mm  $\rightarrow$  117.5 mm)

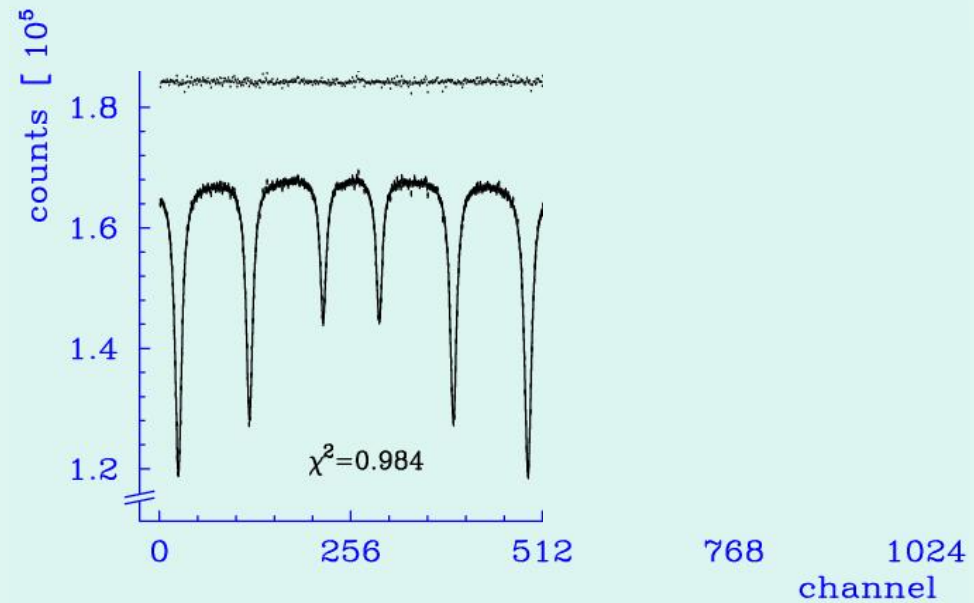
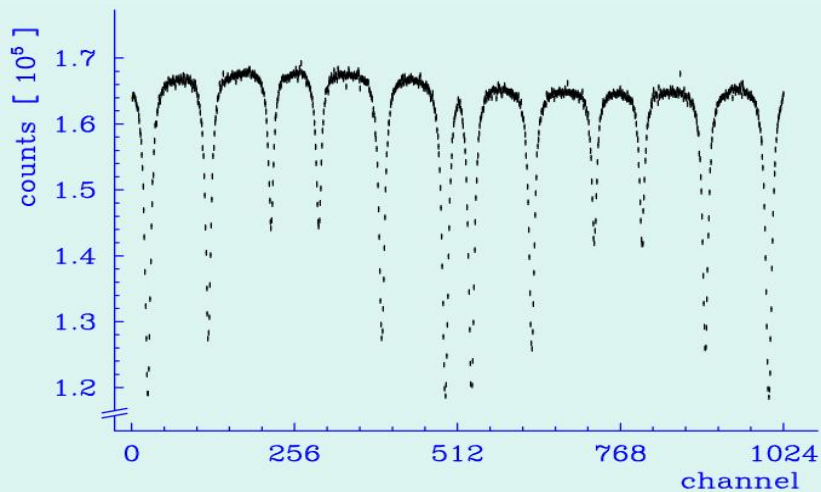
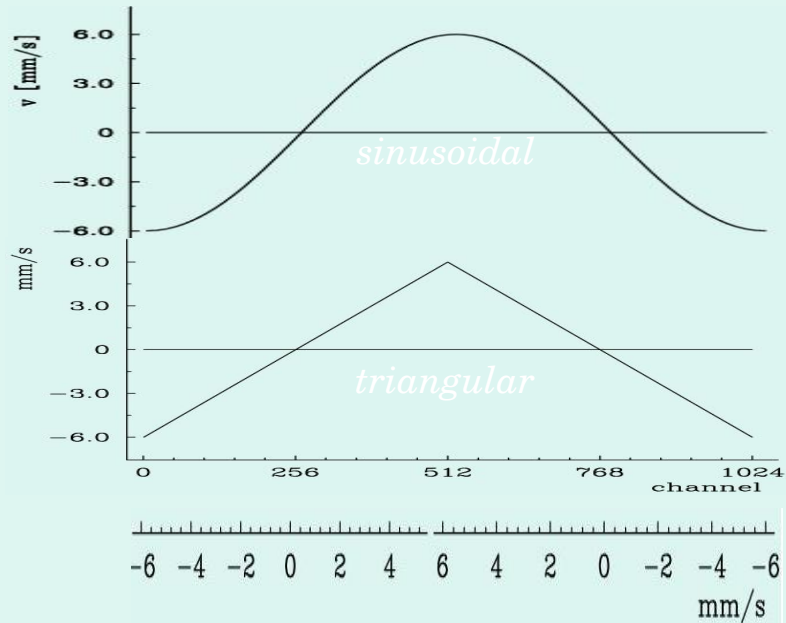
# The raw data problem



*Calibration spectrum for FAS  
W. C. Tennant, Christchurch (2009).*

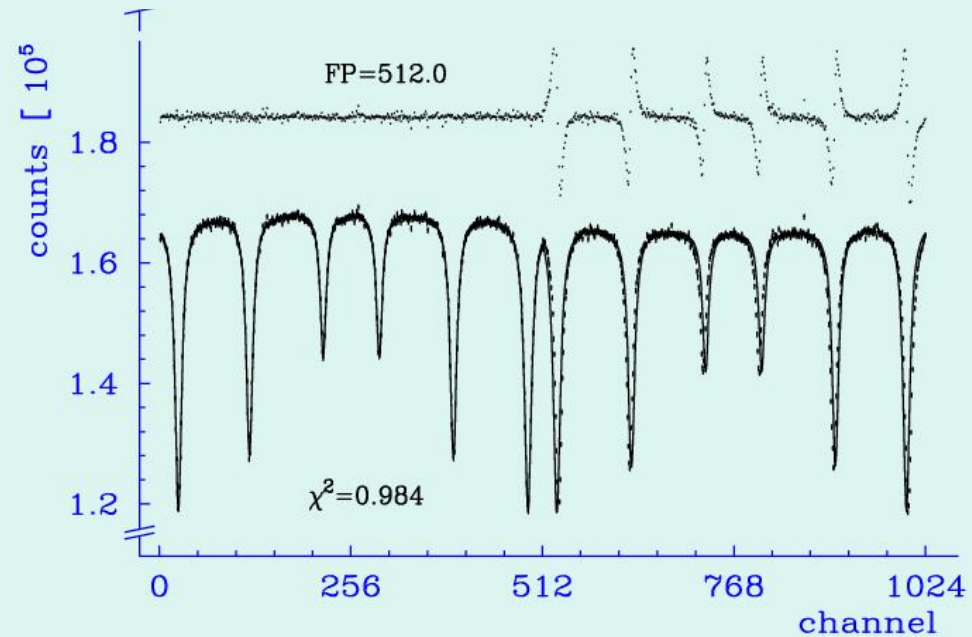
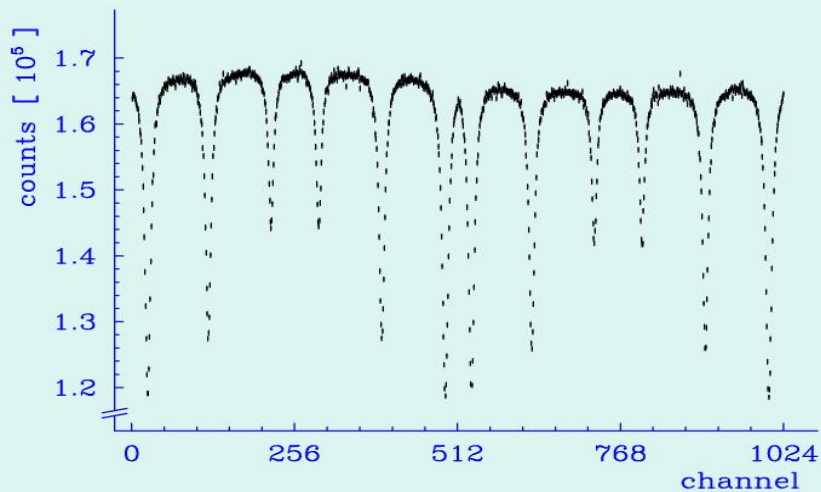
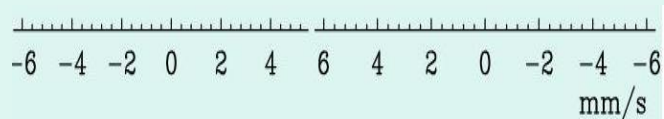
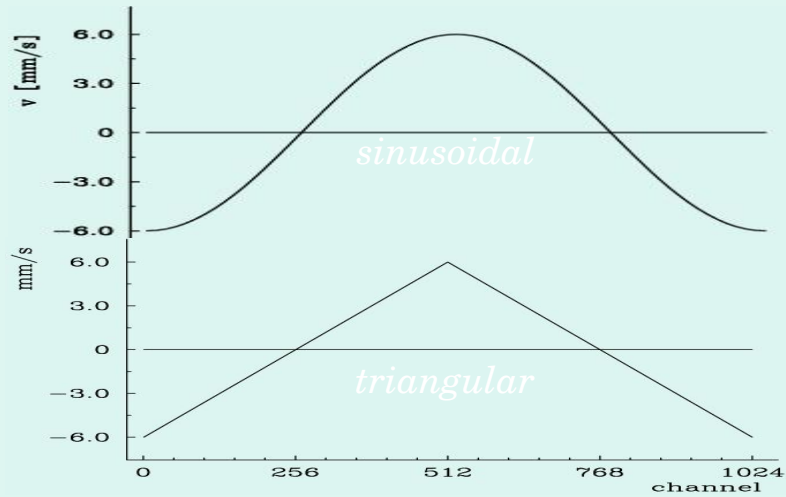
# The raw data problem

*Fit of 6 independent lines  
(position, Voigt profile)*



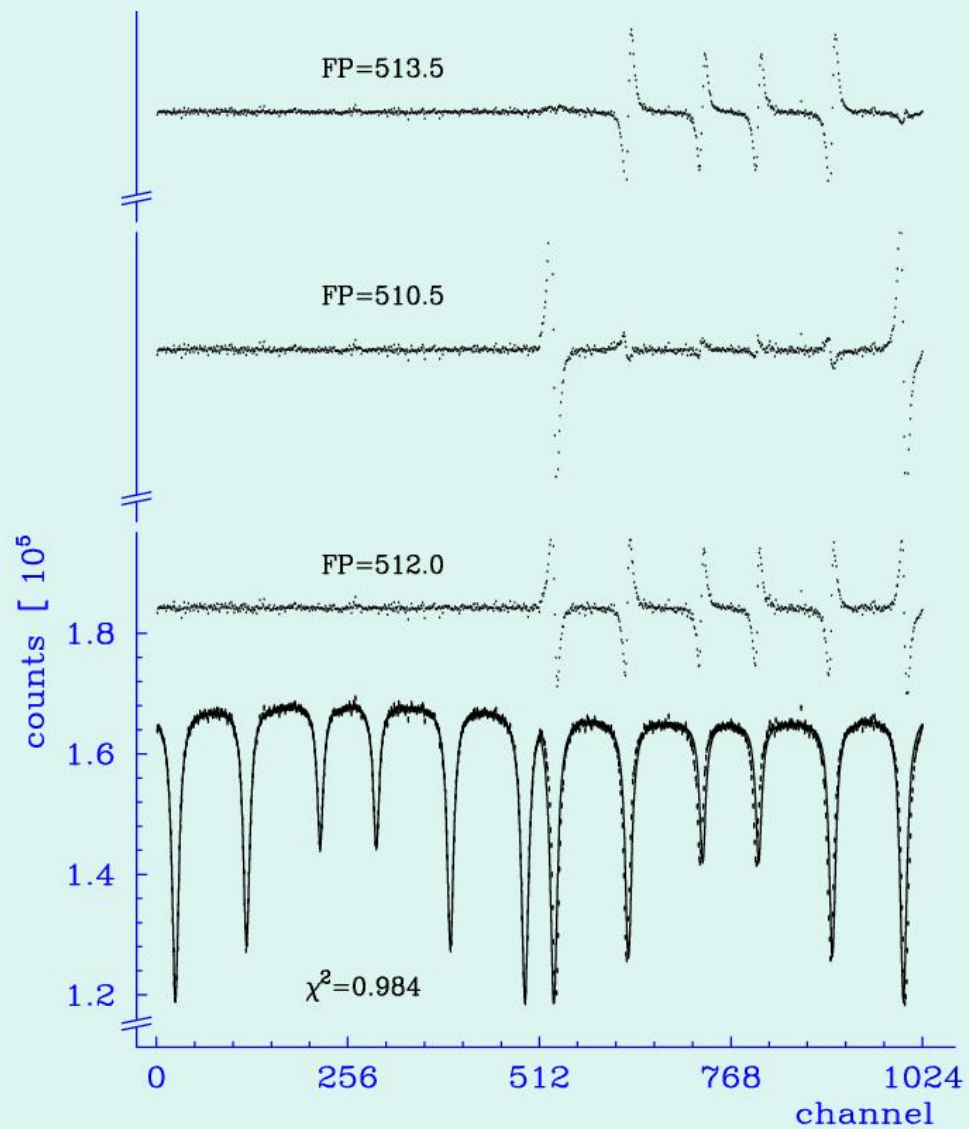
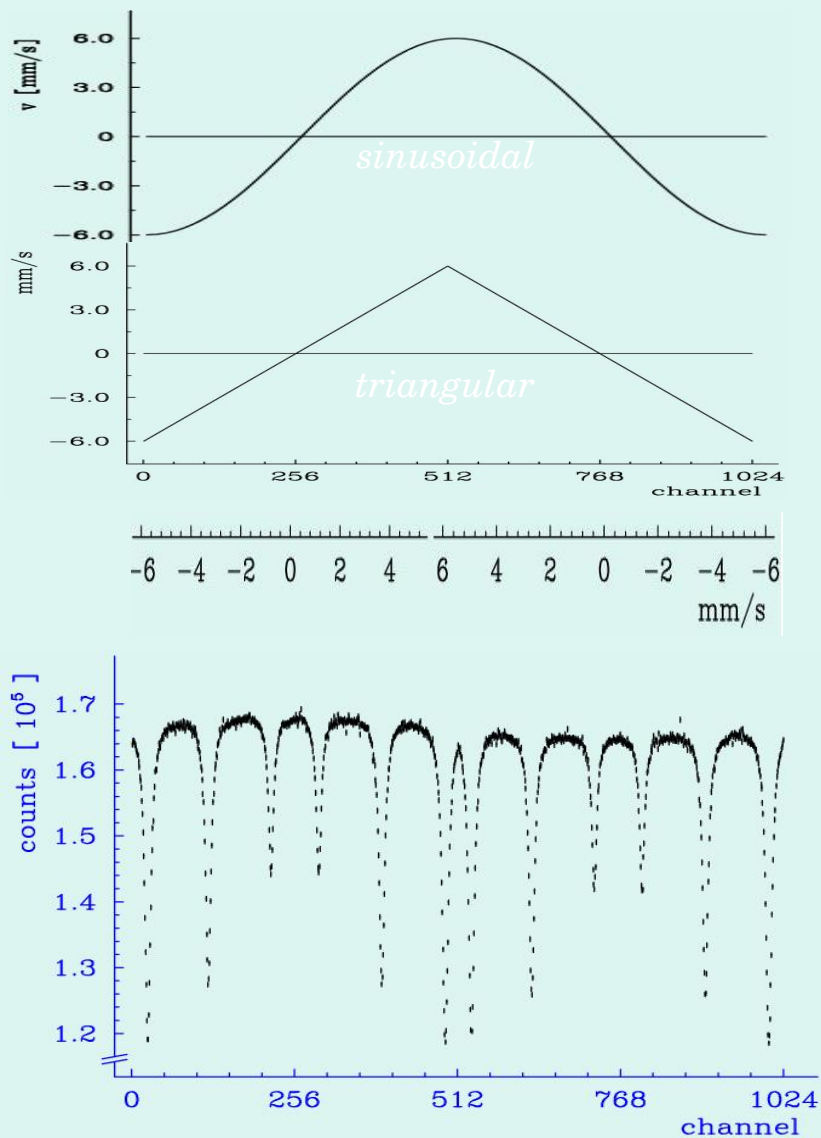
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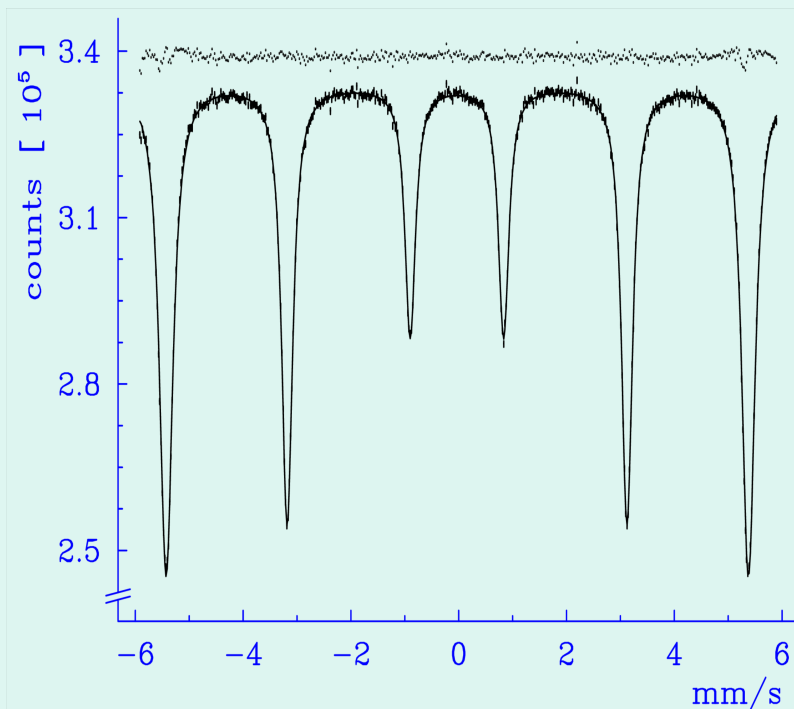


# The raw data problem

*Folding point FP: minimum of  $Z(FP)$*

$$Z(FP) = \sum_{i>FP}^N (C_i - C_{FP-(i-FP)})^2$$

*FP=512.5*



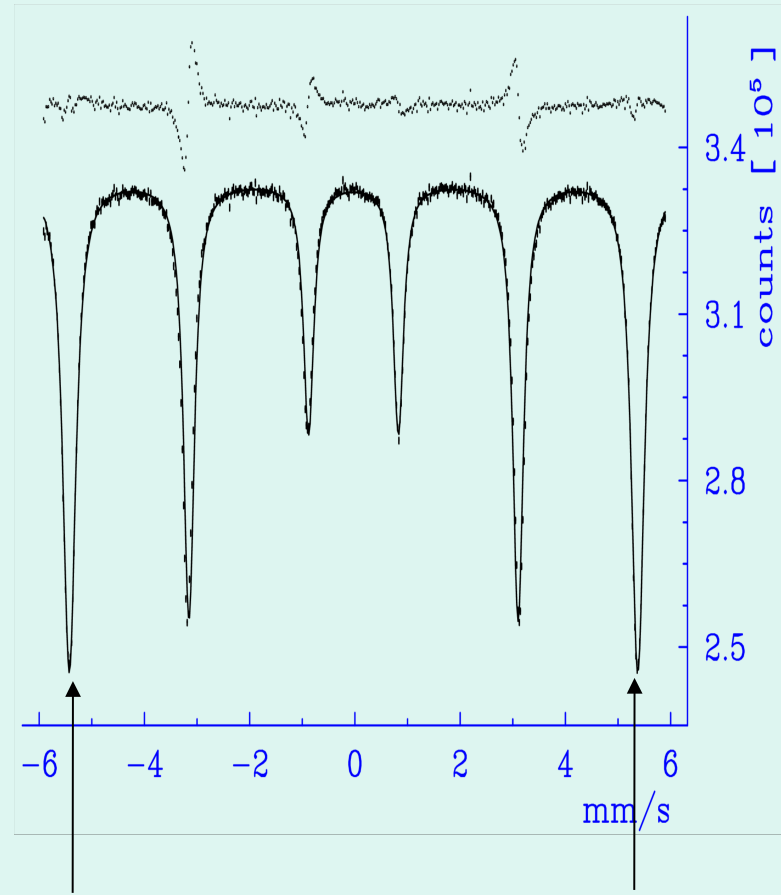
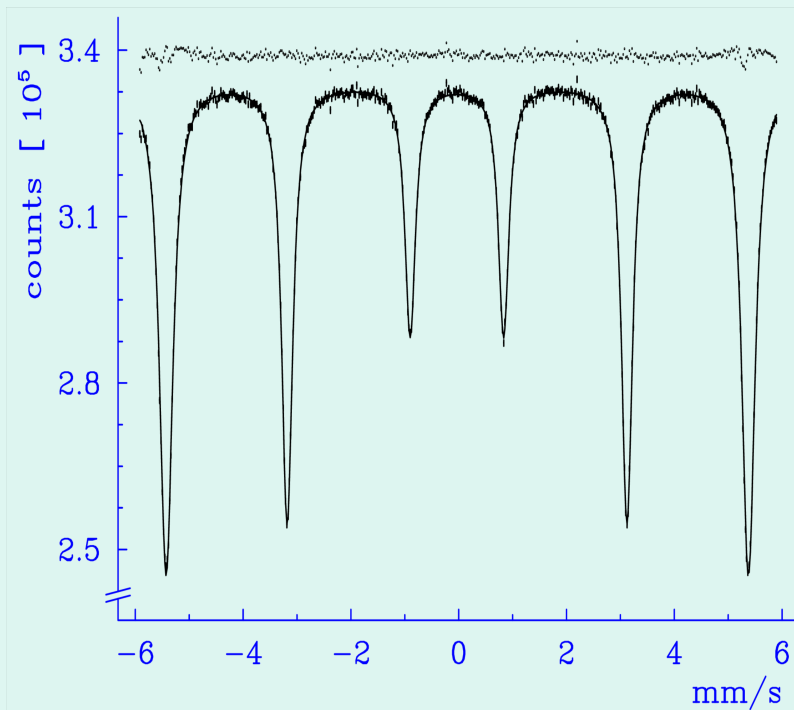
*6 Voigt profiles*     $\chi^2=1.61$

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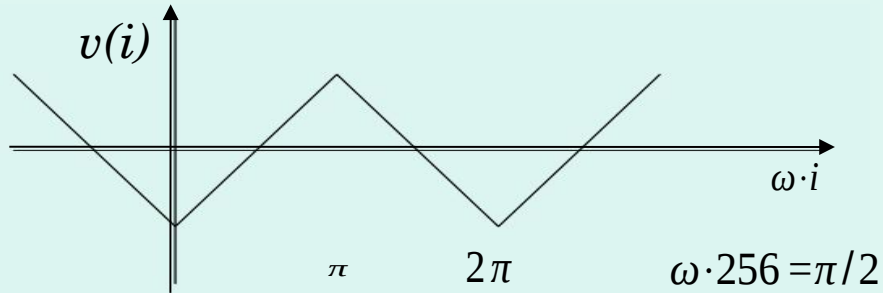
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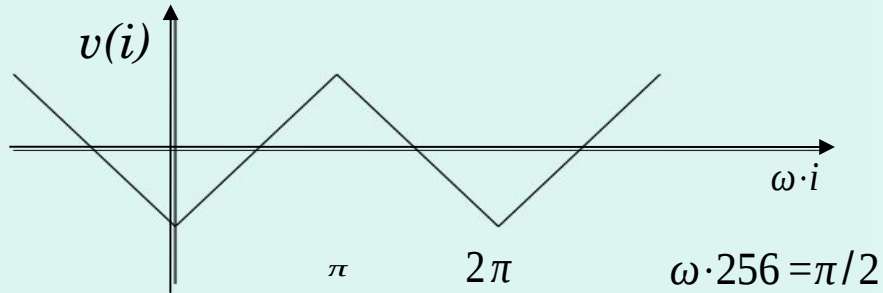
**Correction of the nonlinearity of the velocity scale**

$$v(i) = -v_{max} \frac{4}{\pi^2} \left( \cos \omega \cdot i + \frac{\cos 3 \omega \cdot i}{3^2} + \frac{\cos 5 \omega \cdot i}{5^2} + \dots \right)$$

$$dv(i) = \sum a_k \cos((2k+1)\omega \cdot i) + b_k \sin((2k+1)\omega \cdot i)$$



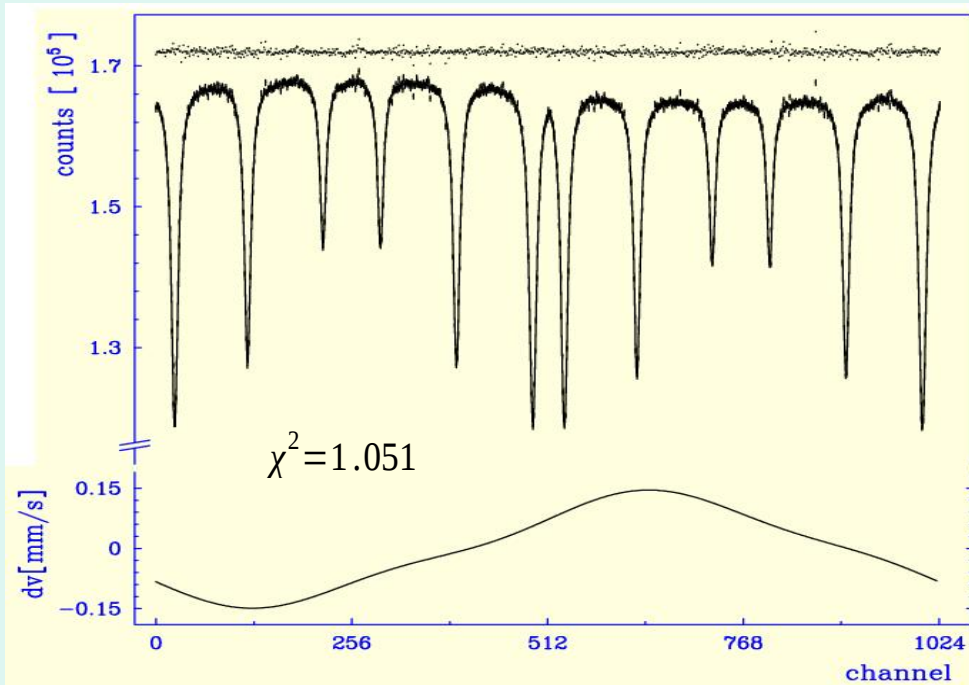
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**Source (5mCi):**

$$\Gamma_{Lorentz} = 1.03 \Gamma_N, \sigma_{Gauss} = 0.45 \Gamma_N$$

$dv = \pm 7.5$  channels

$$\Gamma(\alpha\text{-iron}) = \Gamma_N$$

# ***The raw data problem***

## **Note:**

*No reliable velocity scale after folding!!*

*Better solution by a fit of the full spectrum and velocity correction for each channel*

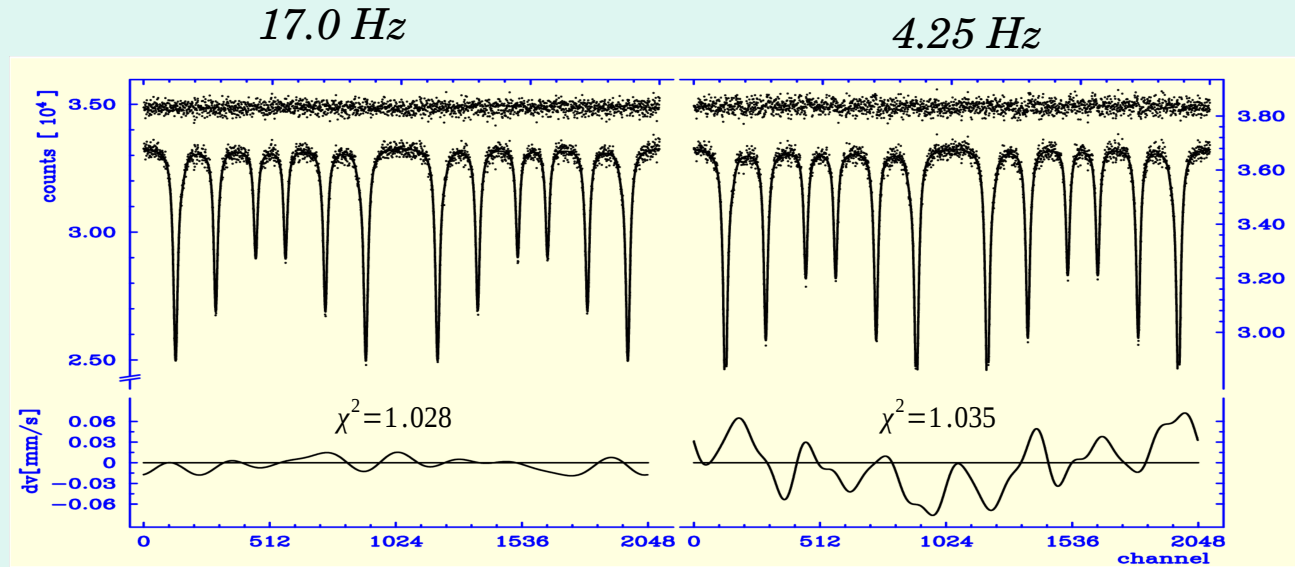
***Not only positions of the lines but also their shapes determine the velocity scale***

# Choice of the drive frequency

J Pechousek et al., Palacky University of Olomouc, Czech Republic

[www.researchgate.net/publication/252974211](http://www.researchgate.net/publication/252974211)

triangular



$$ch_{v=0} = 513.08$$

$$V_{max} = 7.24 \text{ mm/s}$$

$$L_0 = 117.5 \text{ mm}$$

$$geo = \frac{V_{max}}{8 L_0 f}$$

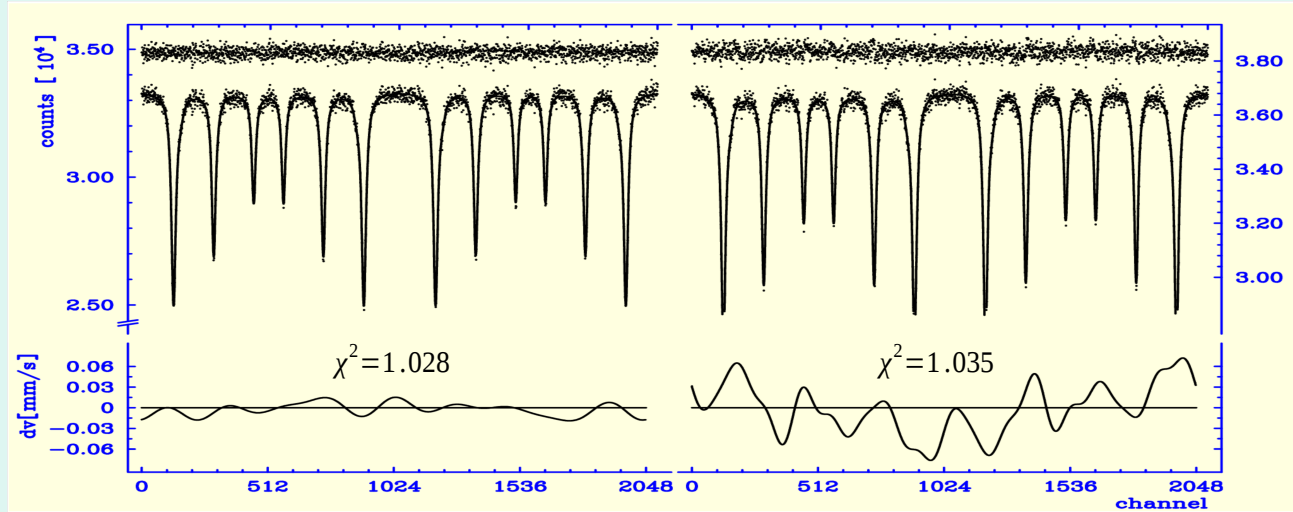
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[www.researchgate.net/publication/252974211](http://www.researchgate.net/publication/252974211)

17.0 Hz

4.25 Hz

triangular



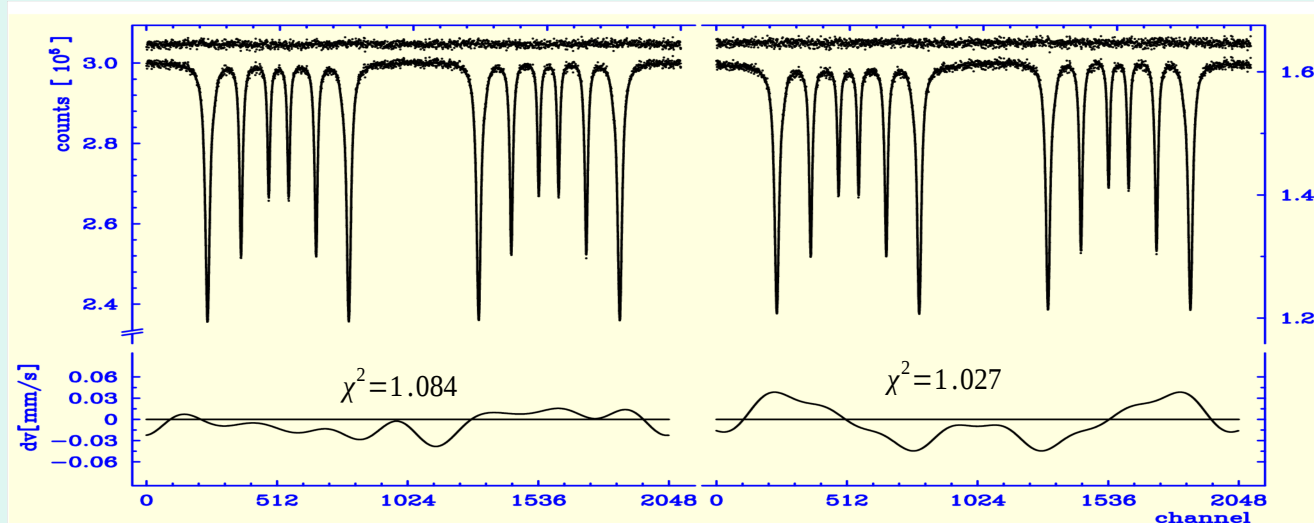
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$$geo = \frac{V_{max}}{8 L_0 f}$$

sinusiodal



$$geo = \frac{V_{max}}{2 \pi L_0 f}$$

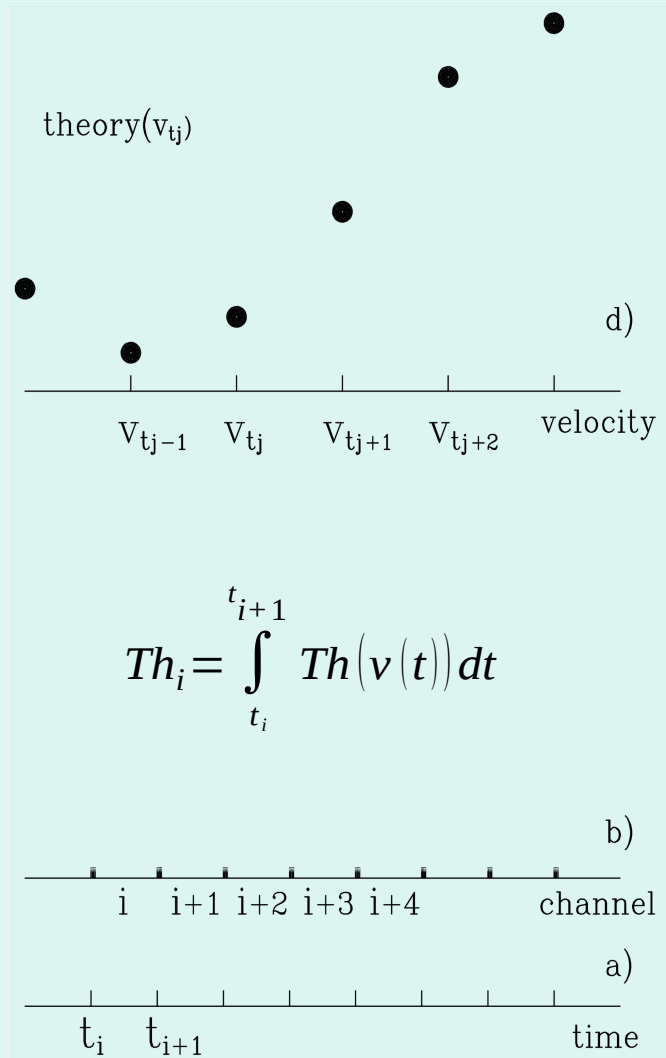
# **Choice of the drive frequency**

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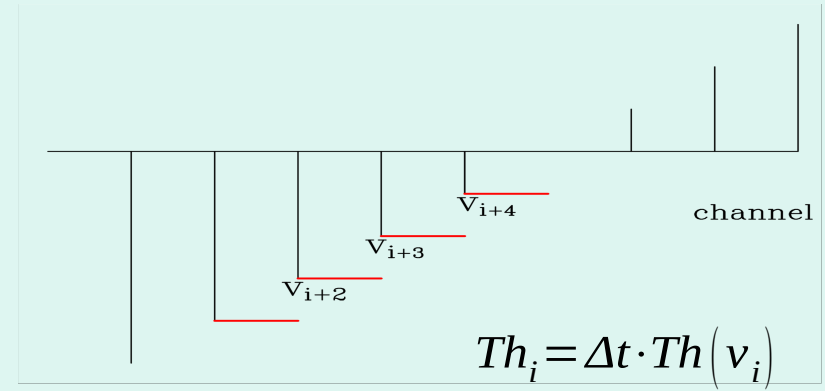
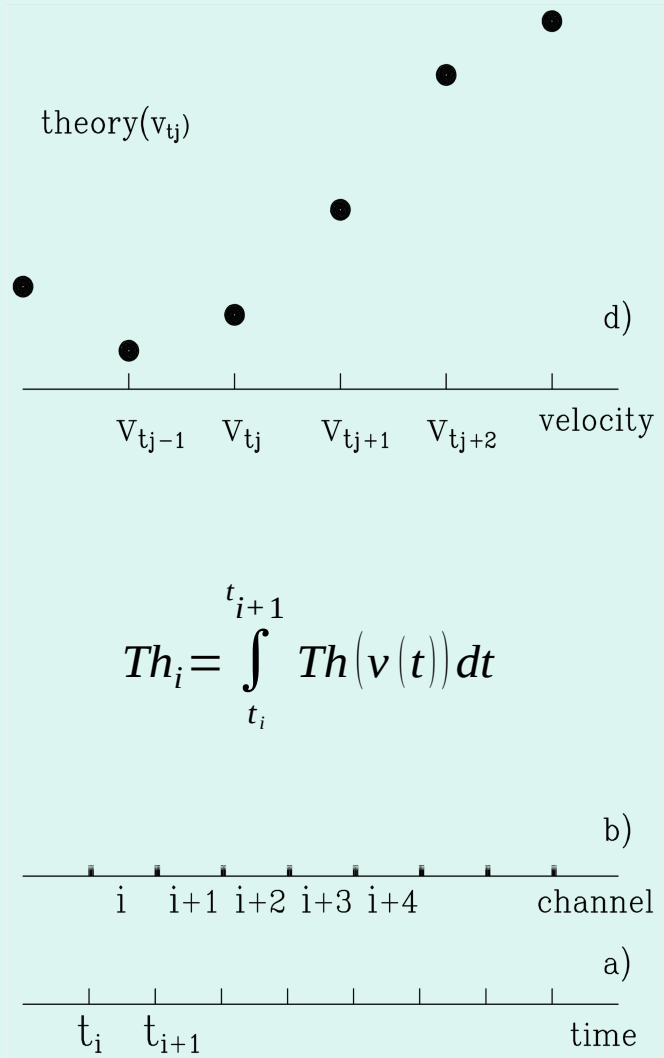
*Nonlinearity correction  $dv(i)$  strongly depends on the drive frequency*

*The evaluated physical parameters of the spectrum shall not depend on the driving mode, the frequency nor the solid angle*

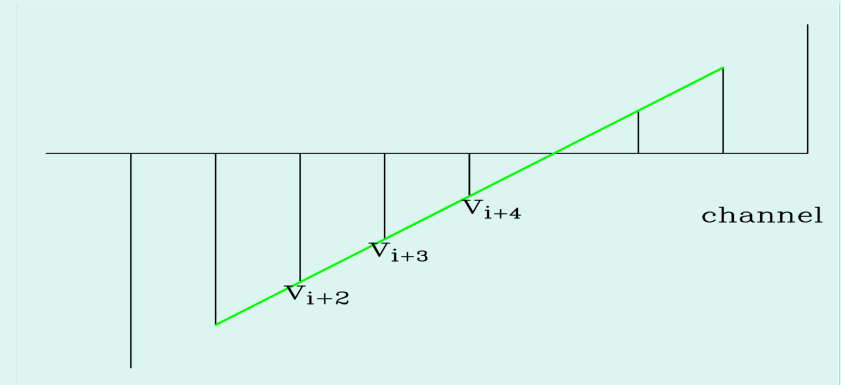
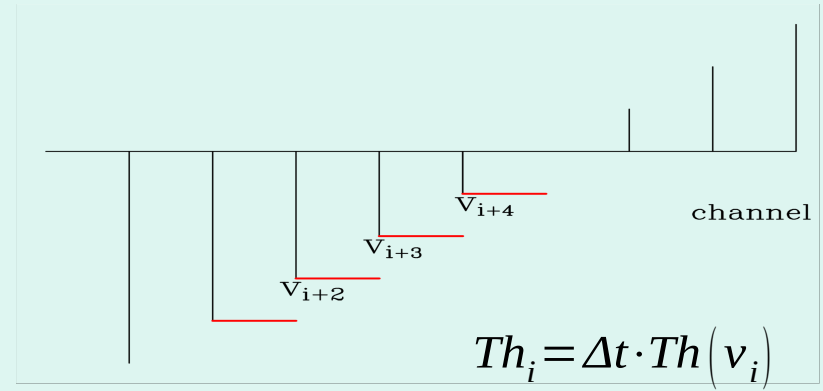
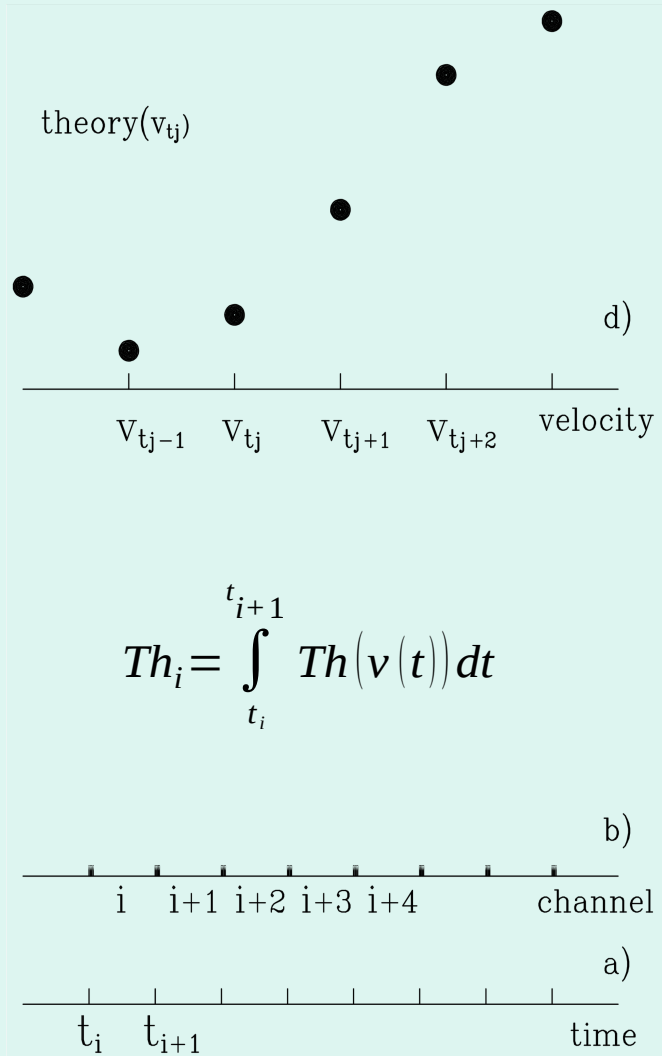
# Theoretical value at channel $i$



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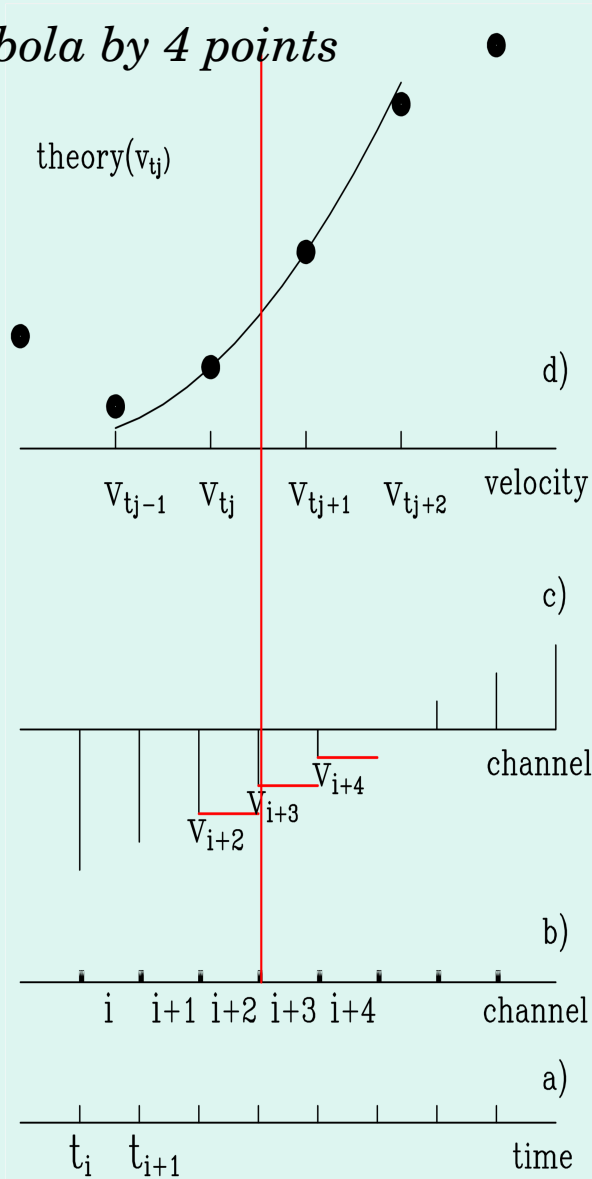
$$Th_i = \int_{t_i}^{t_{i+1}} Th(v(t)) \frac{dt}{dv} dv$$

$$= \frac{\Delta t}{v(t_{i+1}) - v(t_i)} \int_{v(t_i)}^{v(t_{i+1})} Th(v) dv$$



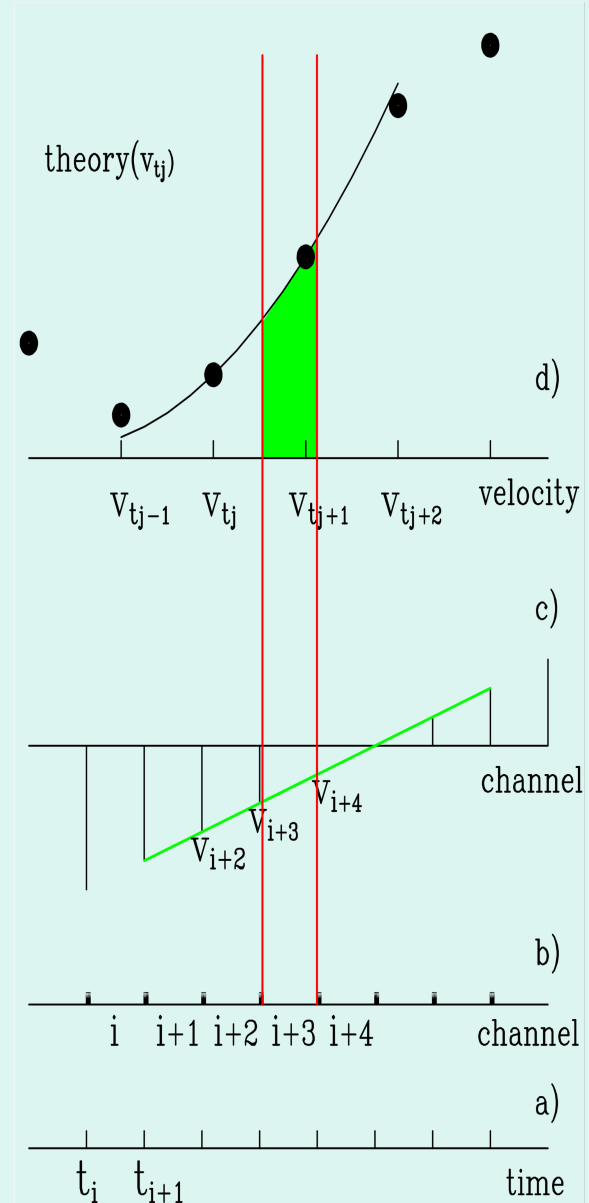
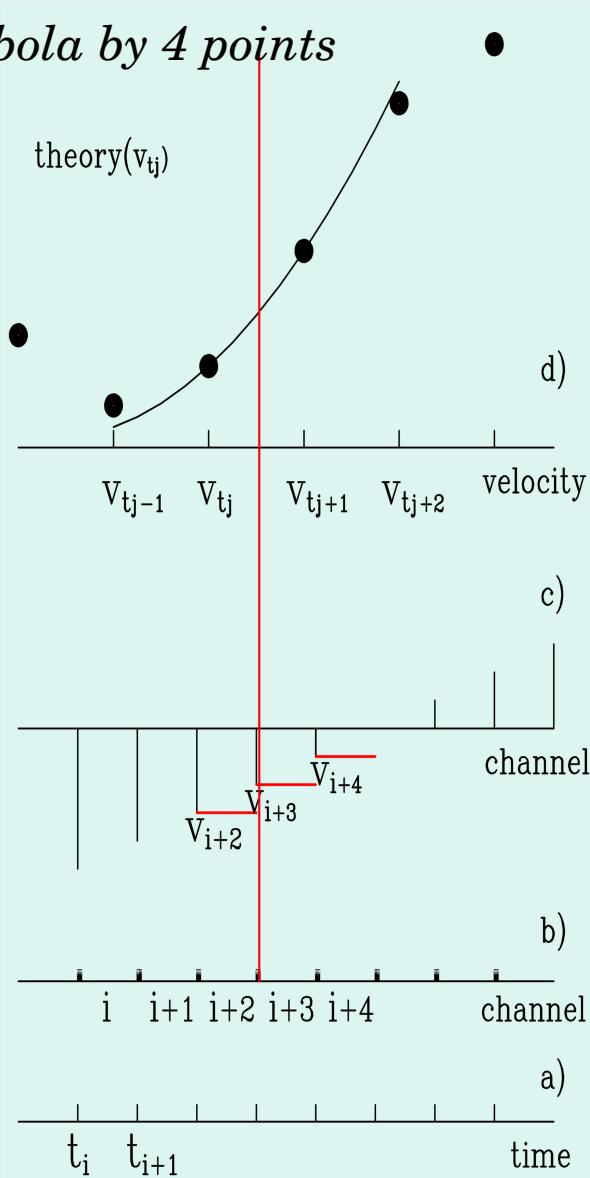
# Theoretical value at channel $i$

*best parabola by 4 points*



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## ***Theoretical value at channel $i$***

### **Note:**

#### ***Advantages of theory interpolation:***

*Fit of  $dv_i$  dependent on 6-18 parameters would not be feasible  
with the convolution for each iteration step: 1024-4096 numerical integrals*

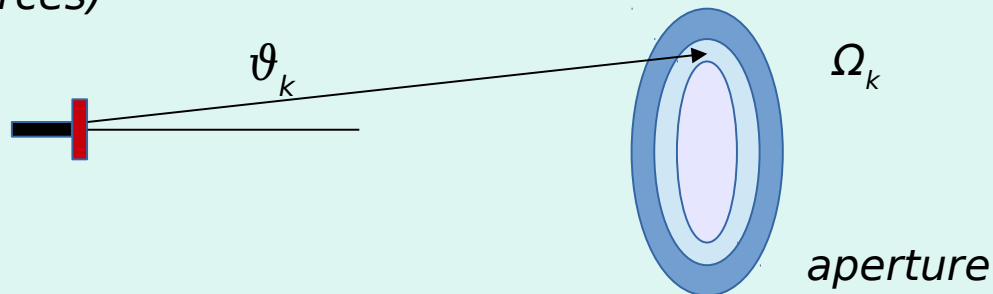
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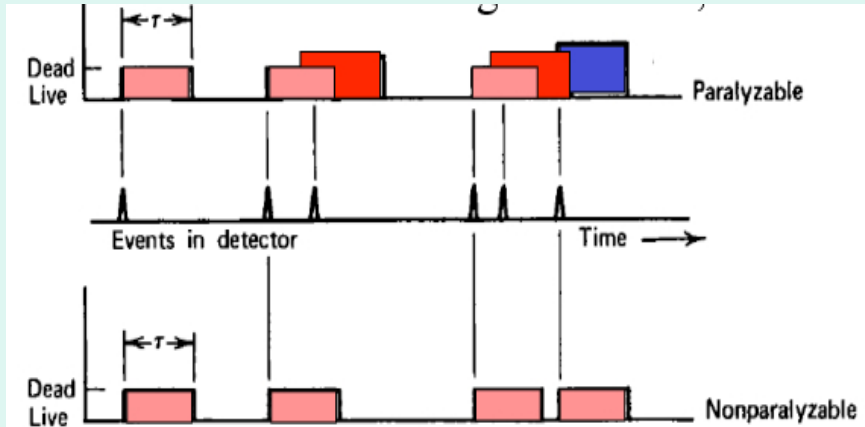
Calculation of cos-smearing for large  $\Omega$  can be easily added (weak sources)



Weighted (by  $\Omega_k$ ) superposition of spectra with  $v_i \cos(\vartheta_k)$

# Dead time

*paralyzable, nonparalyzable*

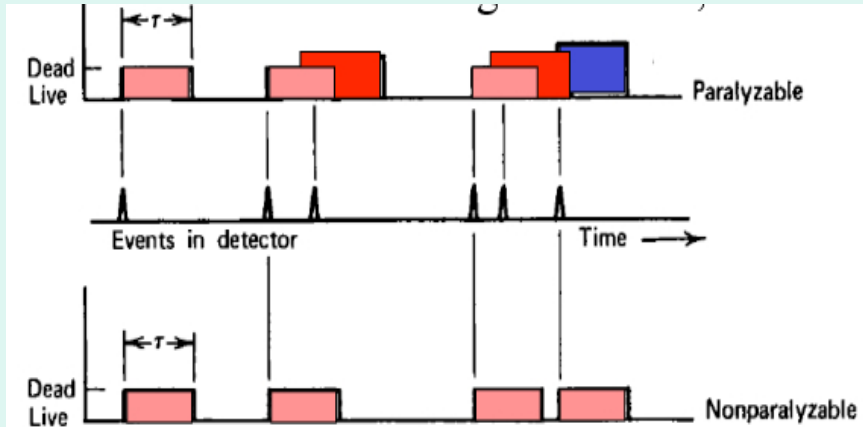


*Proportional counter*  $\tau = 100 \mu\text{s}$

*Scintillator*  $\tau = 1 \mu\text{s}$

# Dead time

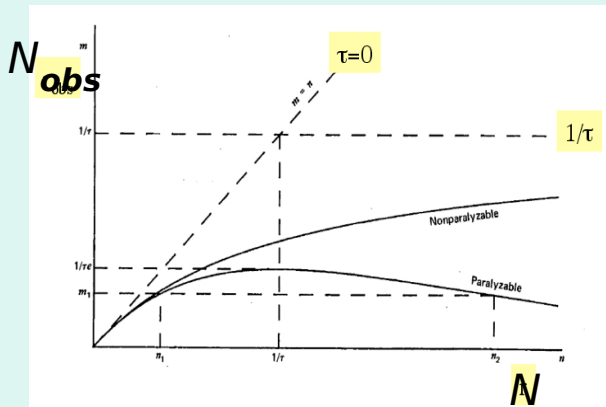
*paralyzable, nonparalyzable*



*Proportional counter*      $\tau = 100 \mu\text{s}$

*Scintillator*

$\tau = 1 \mu\text{s}$



$$N_{obs} = \frac{N}{1 + N\tau_{np}}$$

$$N_{obs} = Ne^{-N\tau_p}$$

$$N_{obs} = \frac{Ne^{-N\tau_p}}{1 + N\tau_{np}}$$

# Dead time

Code for Monte Carlo simulations (nonparalyzable, *paralyzable*)

```
while(isimul[0] < *counts)
{
    eventtime=-log(1.-frand()) / (countrate*theo[ichannel]);
    lastevent=lastevent+eventtime;

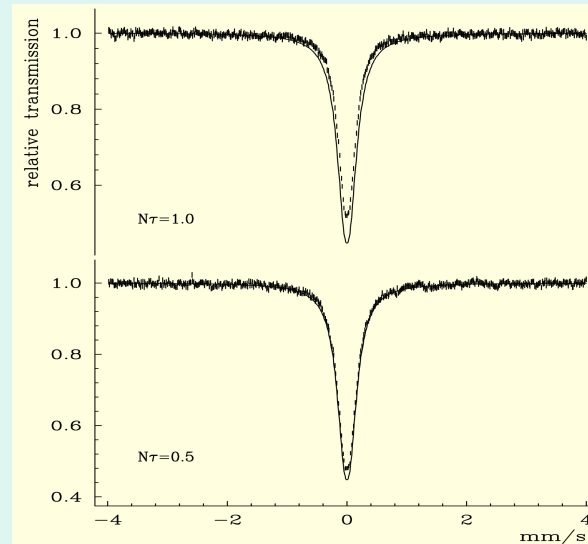
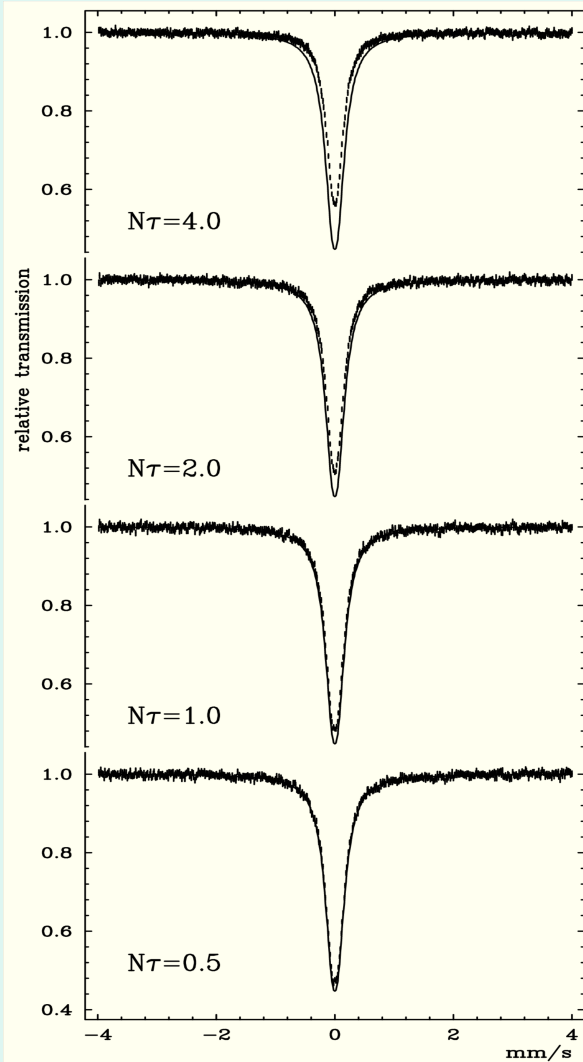
    totaltime=totaltime+eventtime;
    ichannel= totaltime / dwelltime;
    i=ichannel / nu_channel; ichannel=ichannel-i*nu_channel;

    if(lastevent > deadttime)
        {isimul[ichannel]++; lastevent=0.0; icount++;}
    else
        {iloss++;}
        {iloss++;lastevent=0.;}
}
```

# Dead time

*Nonparalyzable*

*paralyzable*

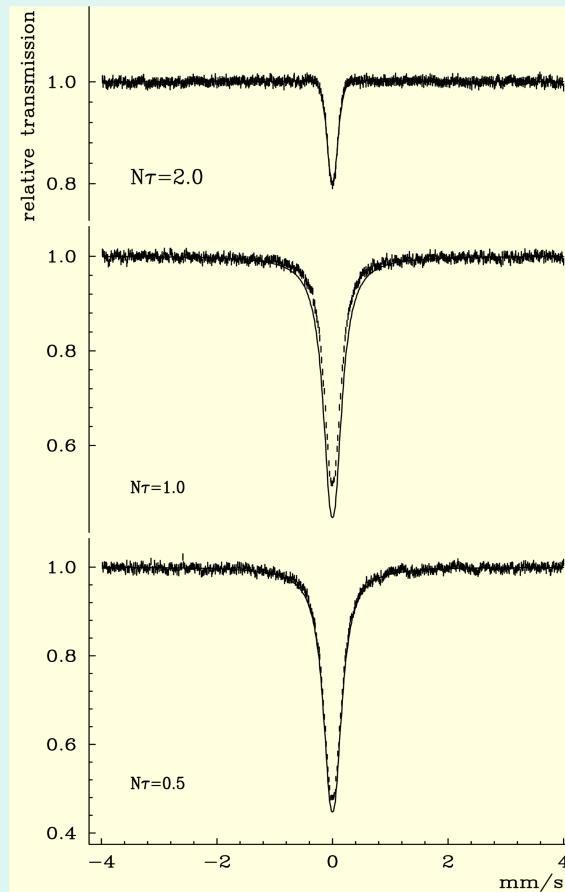
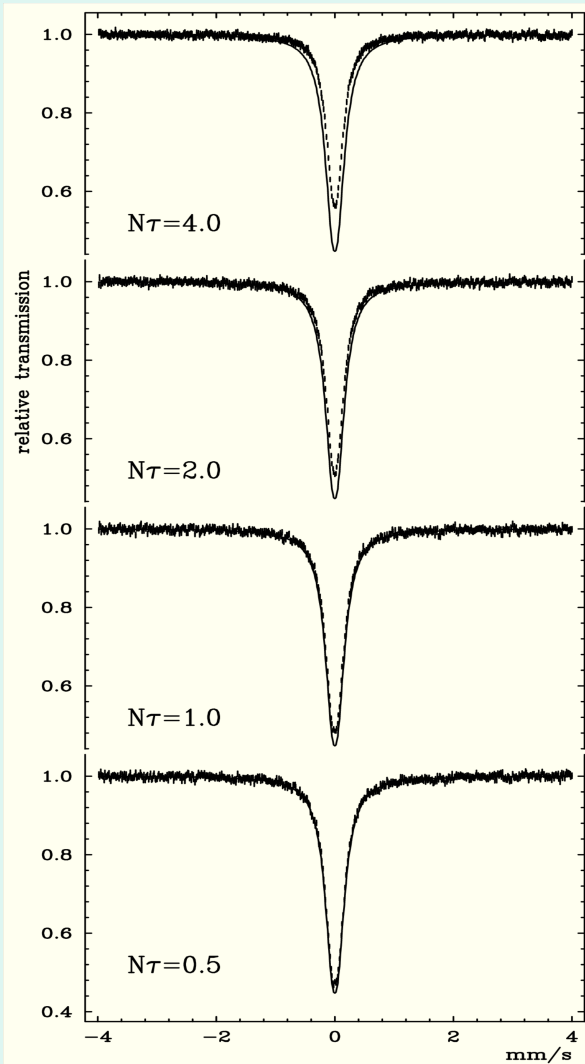




# Dead time

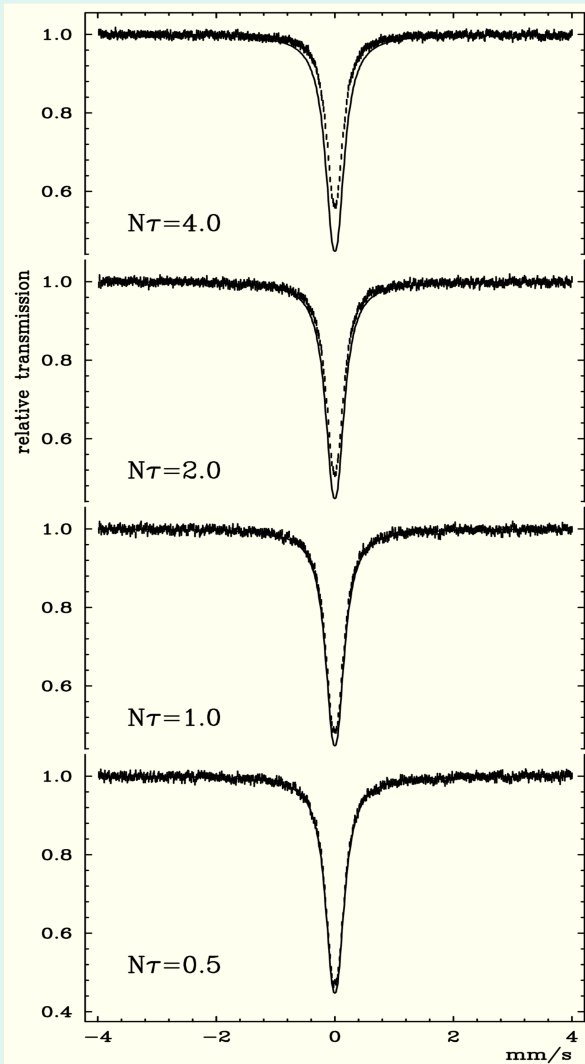
*Nonparalyzable*

*paralyzable*

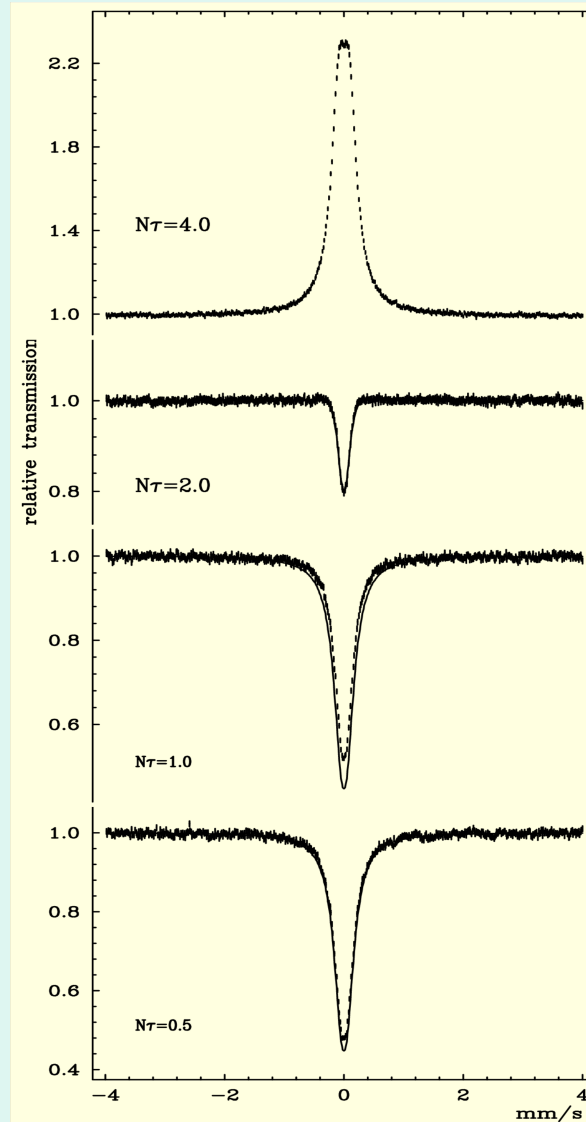


# Dead time

*Nonparalyzable*

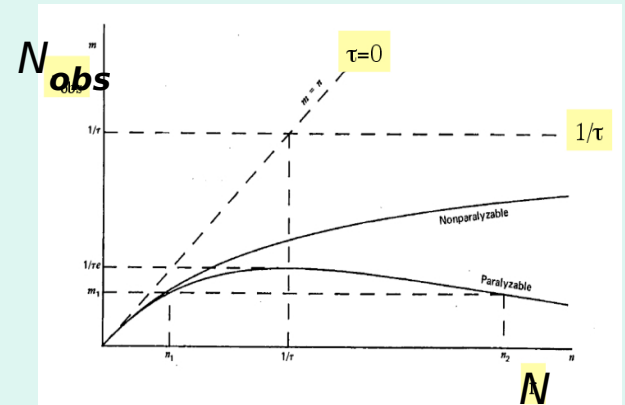
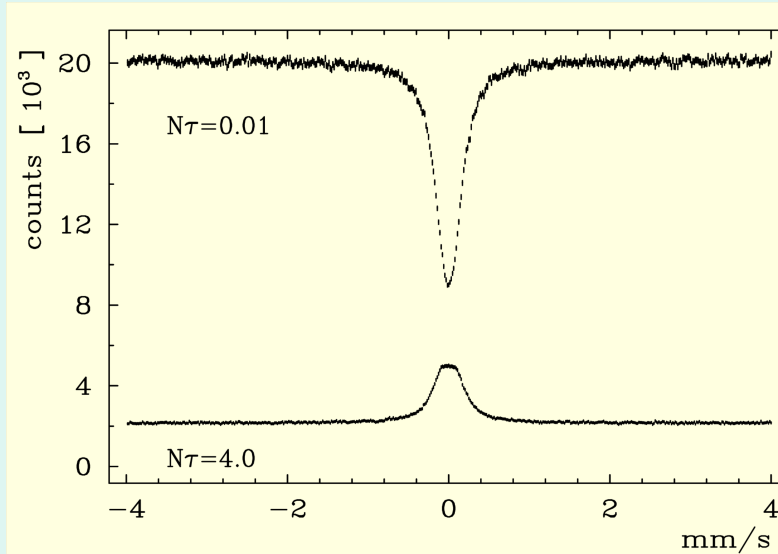


*paralyzable*



# Dead time

*paralyzable*

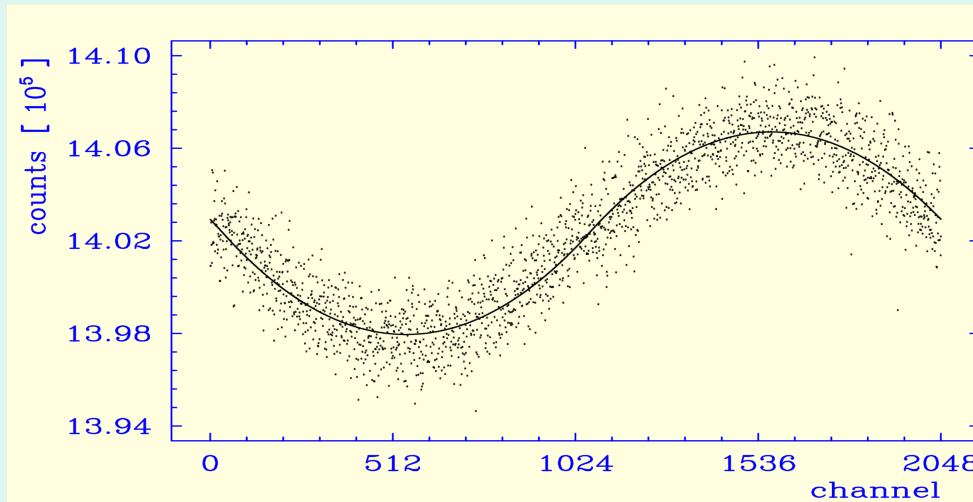


**up to  $N\tau \leq 0.5$  deadtime effects are kept within reasonable limits**

# Dead time

$$geo = \frac{v_{max}}{2\pi L_0 f} = \frac{15.6 \text{ mm/s}}{2\pi \cdot 67 \text{ mm} \cdot 16 \text{ Hz}} = 2.32 \cdot 10^{-3}$$

**5.5mCi**



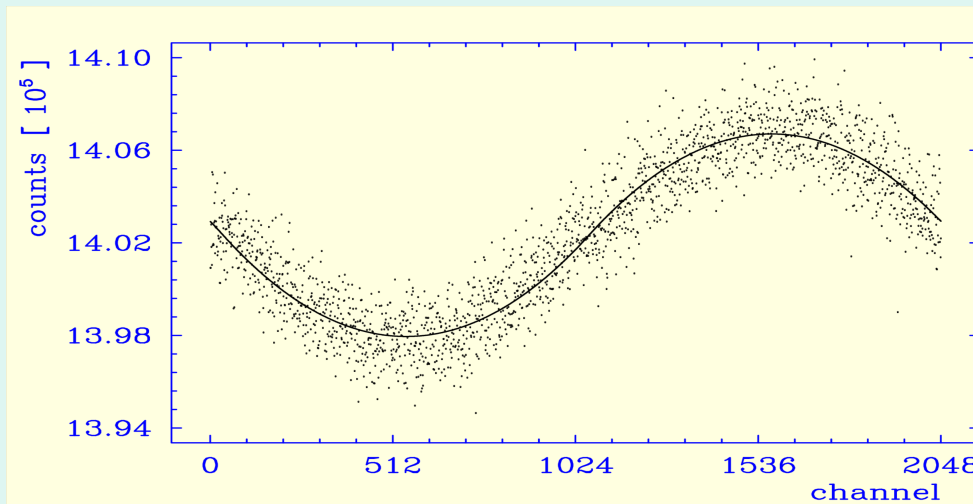
$$\chi^2 = 0.994$$

$$N = counts(v_\infty) \frac{\Omega(geo, v_i)}{\Omega_0} \quad \Rightarrow \quad geo = 1.62 \cdot 10^{-3} \quad (fit!)$$

# Dead time

$$geo = \frac{v_{max}}{2\pi L_0 f} = \frac{15.6 \text{ mm/s}}{2\pi \cdot 67 \text{ mm} \cdot 16 \text{ Hz}} = 2.32 \cdot 10^{-3}$$

**5.5mCi**



$$\chi^2 = 0.994$$

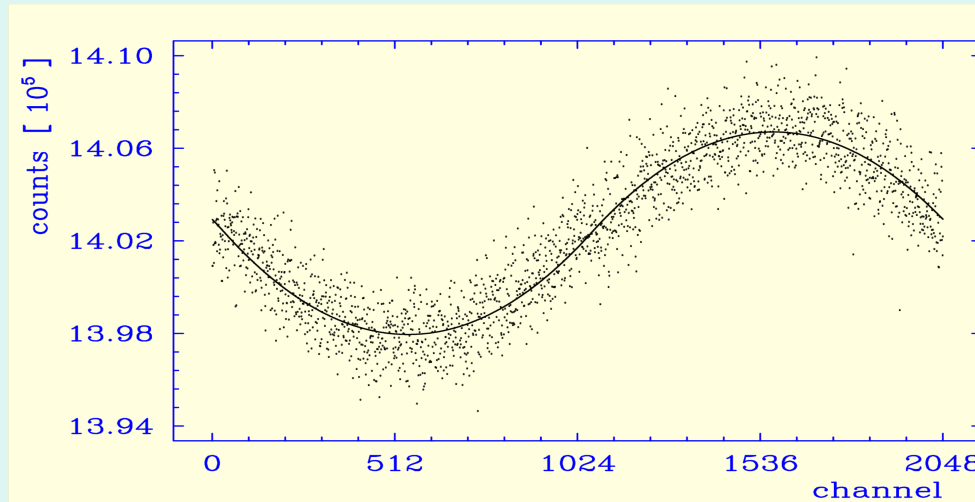
$$N = \text{counts}(v_\infty) \frac{\Omega(geo, v_i)}{\Omega_0} \quad \Rightarrow \quad geo = 1.62 \cdot 10^{-3} \quad (\text{fit!})$$

$$N = \text{counts}(v_\infty) \frac{\Omega(2.32 \cdot 10^{-3}, v_i)}{\Omega_0}, \quad N_c = \frac{N}{1 + N\tau_{np}} \quad \Rightarrow \quad N\tau_{np} = 0.42 \quad (\text{fit!})$$

# Dead time

$$geo = \frac{v_{max}}{2\pi L_0 f} = \frac{15.6 \text{ mm/s}}{2\pi \cdot 67 \text{ mm} \cdot 16 \text{ Hz}} = 2.32 \cdot 10^{-3}$$

**5.5mCi**



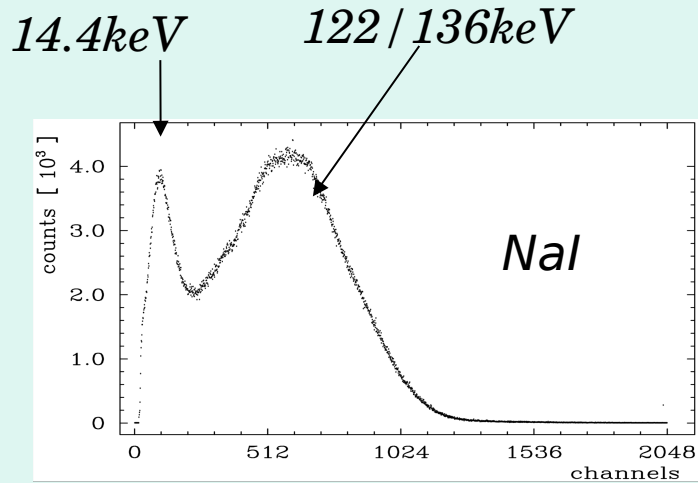
$$\chi^2 = 0.994$$

$$N = \text{counts}(v_\infty) \frac{\Omega(geo, v_i)}{\Omega_0} \Rightarrow geo = 1.62 \cdot 10^{-3} \quad (\text{fit!})$$

$$N = \text{counts}(v_\infty) \frac{\Omega(2.32 \cdot 10^{-3}, v_i)}{\Omega_0}, \quad N_c = \frac{N}{1 + N\tau_{np}} \Rightarrow N\tau_{np} = 0.42 \quad (\text{fit!})$$

$$\frac{2.32}{(1 + N\tau_{np})} = 1.63$$

# Dead time



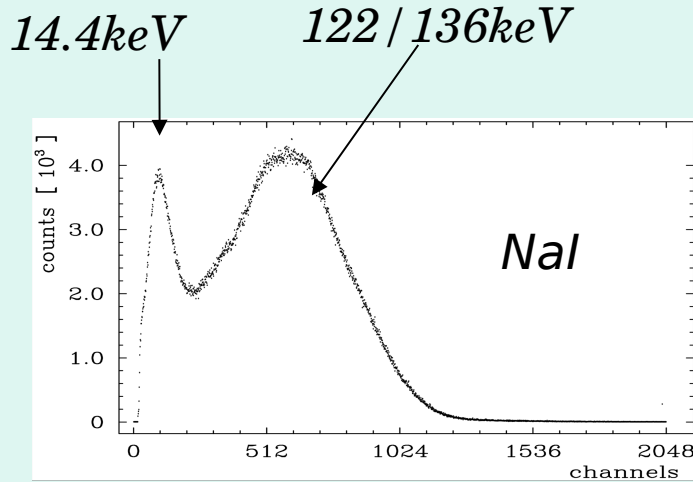
*(ungated)*

*nonparalyzable*

$$N_c = \frac{N}{1 + N\tau_{np}}$$

# Dead time

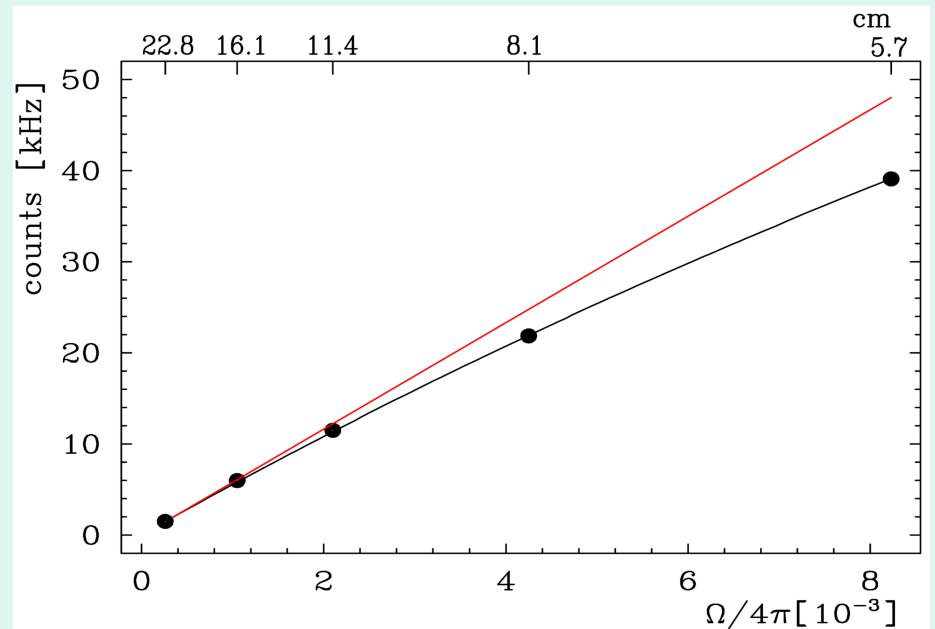
(ungated)



nonparalyzable

$$N_c = \frac{N}{1 + N\tau_{np}}$$

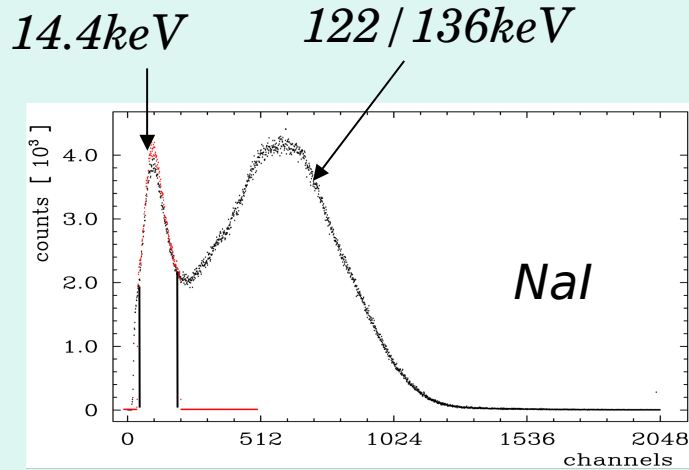
$$N\tau_{np} = 0.22 \rightarrow \tau_{np} = 4.4 \mu\text{s}$$





# Dead time

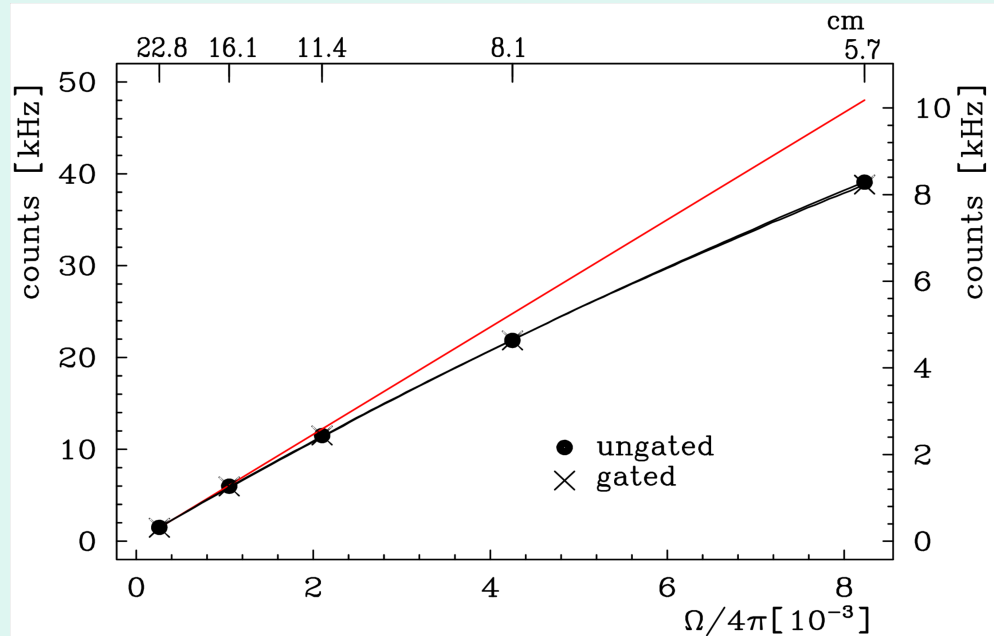
(gated)



nonparalyzable

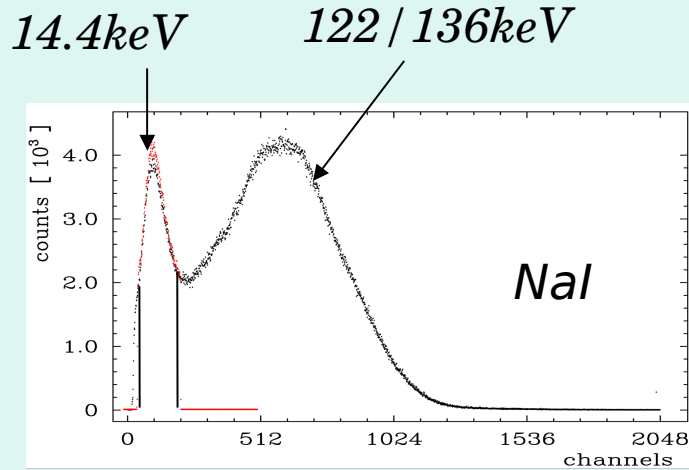
$$N_c = \frac{N}{1 + N\tau_{np}}$$

$$N\tau_{np} = 0.22 \rightarrow \tau_{np} < 4.4 \mu\text{s}$$



# Dead time

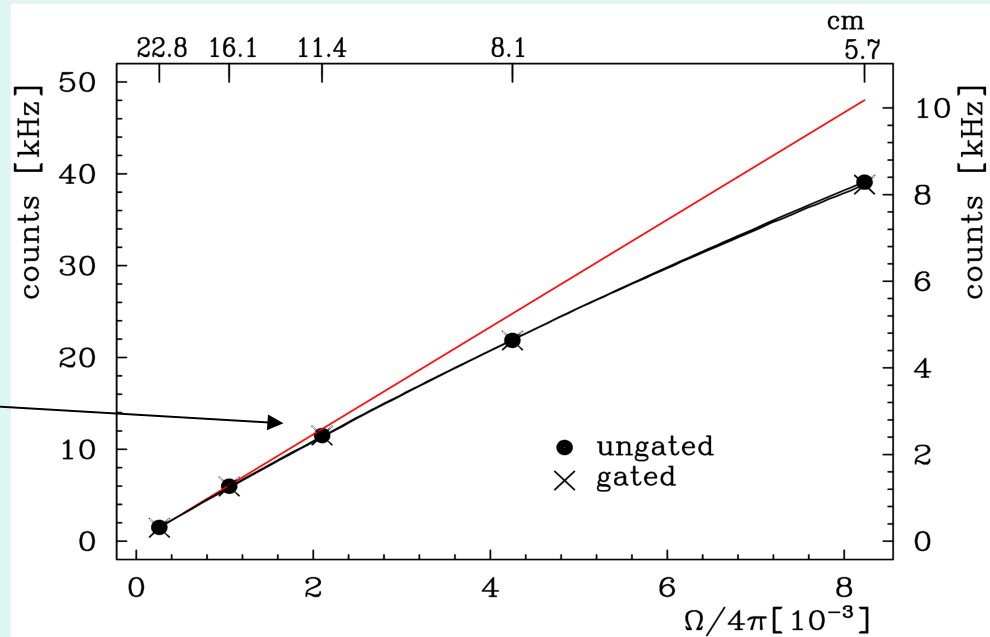
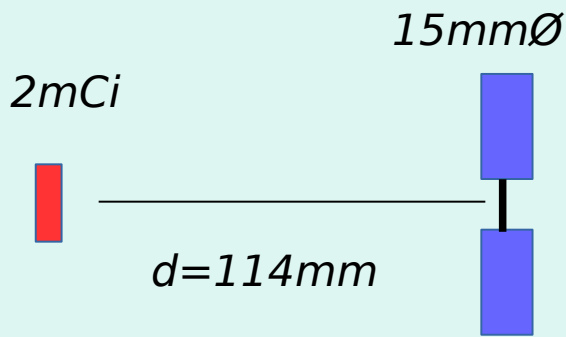
(gated)



nonparalyzable

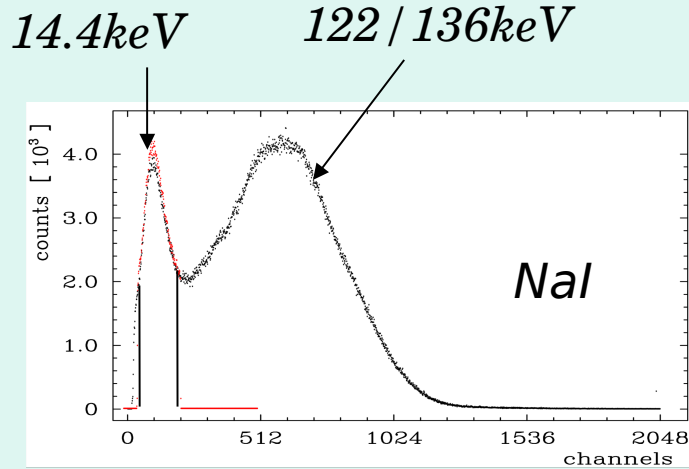
$$N_c = \frac{N}{1 + N\tau_{np}}$$

$$N\tau_{np} = 0.22$$



# Dead time

(gated)



nonparalyzable

$$N_c = \frac{N}{1 + N\tau_{np}}$$

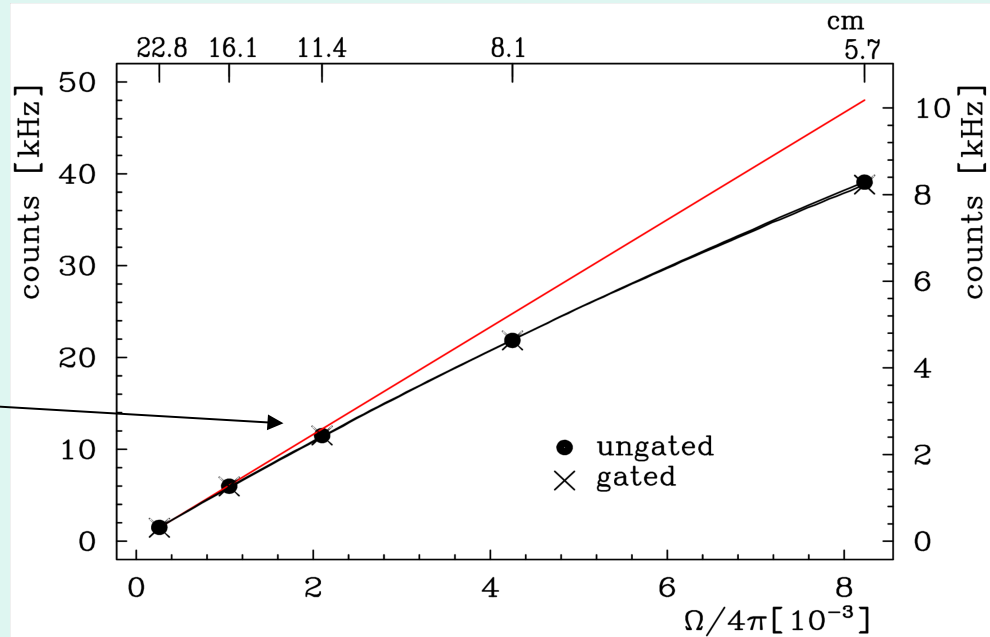
$$N\tau_{np} = 0.22$$

2mCi

15mmØ

d=114mm

50mCi → 5d



# **Dead time**

## **Note:**

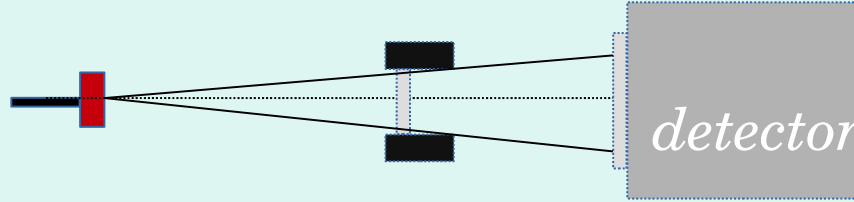
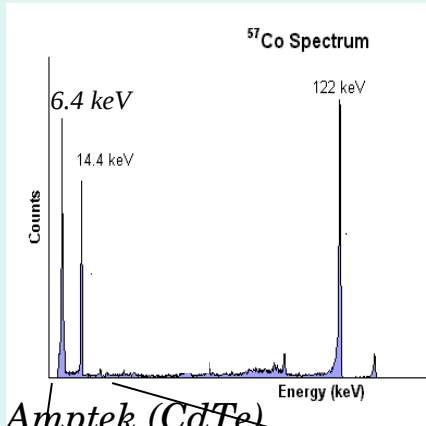
*Avoid (severe) dead time effects !*

*Check by the distance law  $1/r^2$  for dead time*

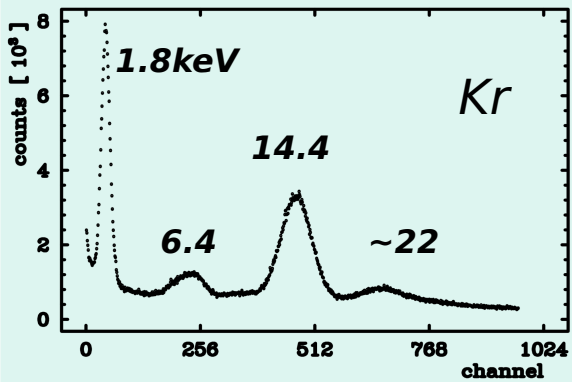
*Check by the geo parameter for dead time:  $\frac{geo}{geo_{eff}} = (1 + N\tau)$*

*up to  $N\tau \leq 0.5$  deadtime effects are kept within reasonable limits*

# Background

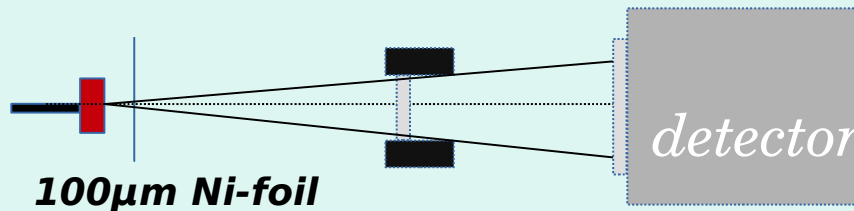
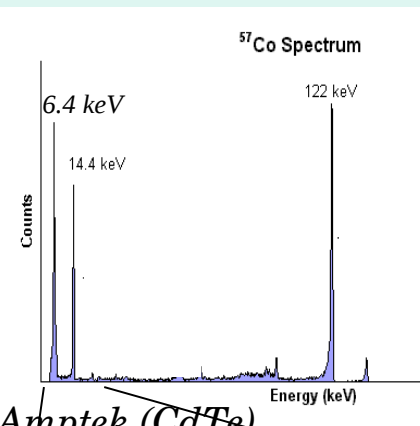


*Amptek (CdTe)*

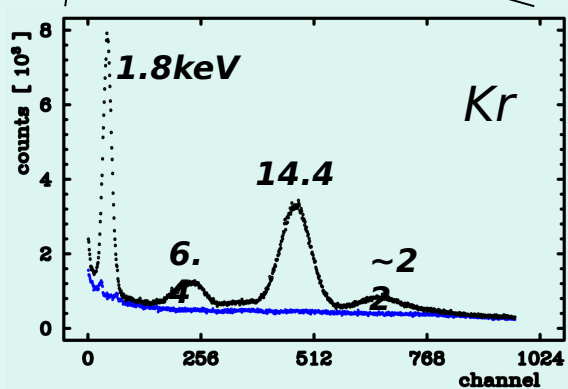


**A straightforward experimental method to evaluate the Lamb-Mössbauer factor of a <sup>57</sup>Co/Rh source**  
G. Spina , M. Lantieri, Nul. Instrum. Methods B 318 (2014) 253-257

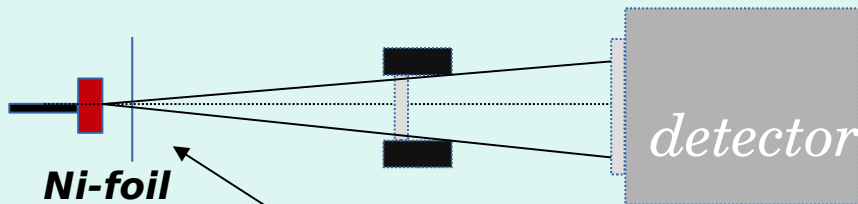
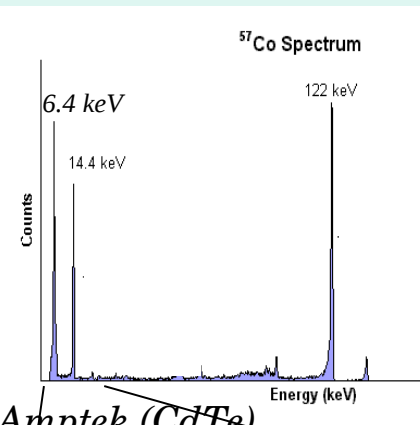
# Background



*Amptek (CdTe)*

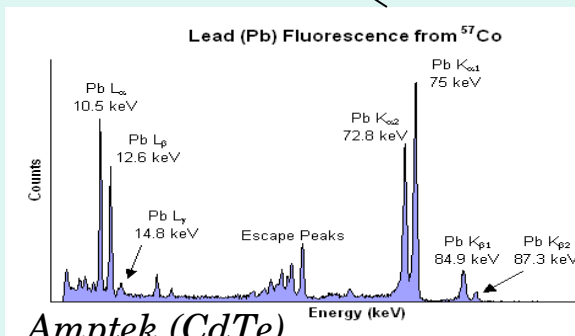


# Background

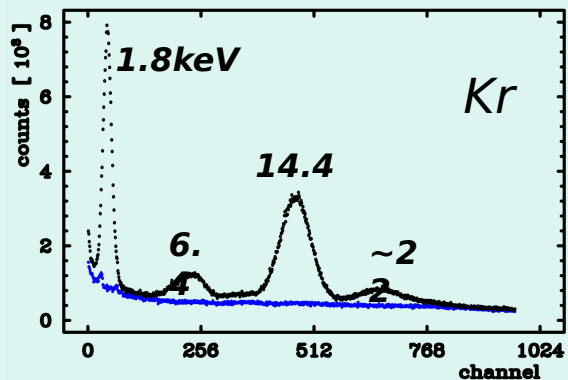


**Position behind source**

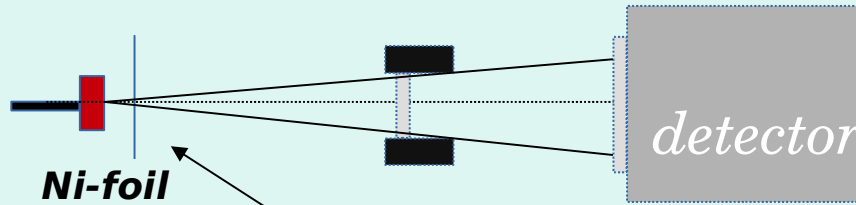
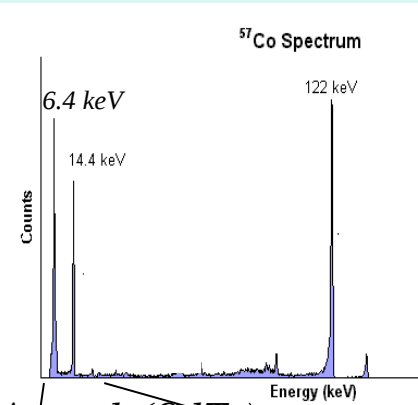
*Amptek (CdTe)*



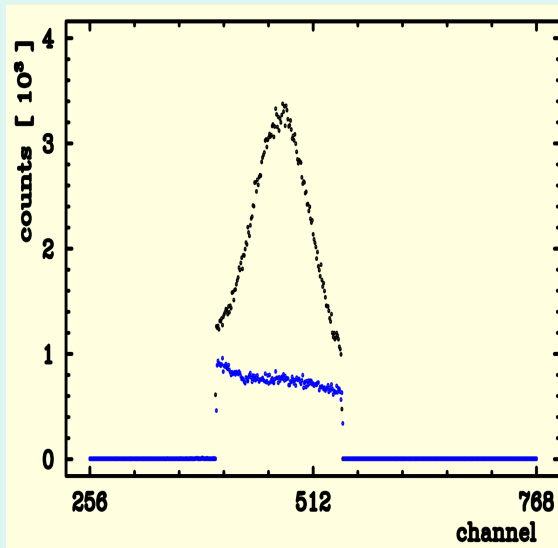
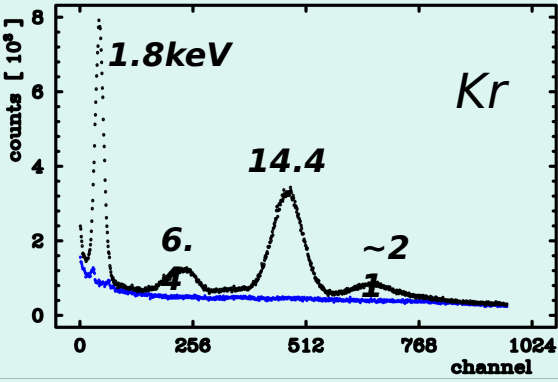
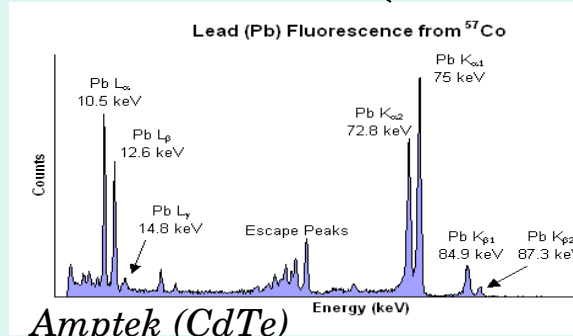
*Amptek (CdTe)*



# Background



*Amptek (CdTe)*



14.4keV - window  
multiscaling mode

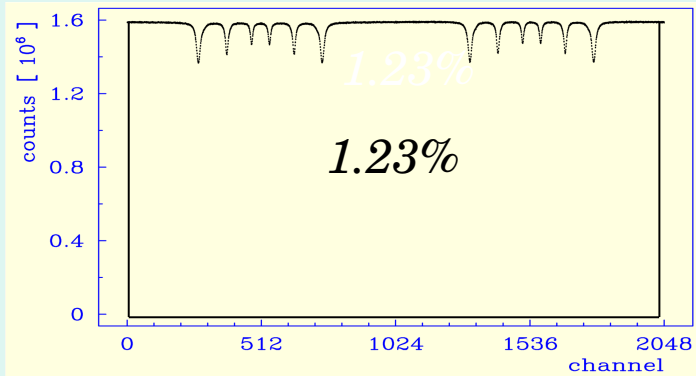
$$A_b(\text{Ni}), A_t$$

Attenuation of 122 / 136 keV  
110 μNi: 0.9695

$$bg_{fr} = \frac{A_b(\text{Ni})/0.9695}{A_t}$$

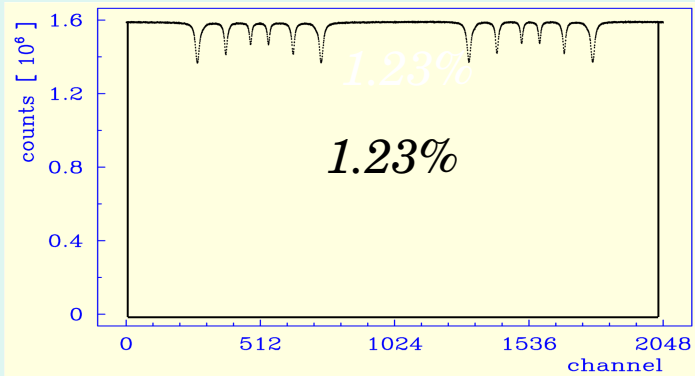


# Background



$$b_{fr} = \frac{A_b(\text{Ni})/0.9695}{A_t/0.9877} = 1.019 \frac{A_b(\text{Ni})}{A_t}$$

# Background

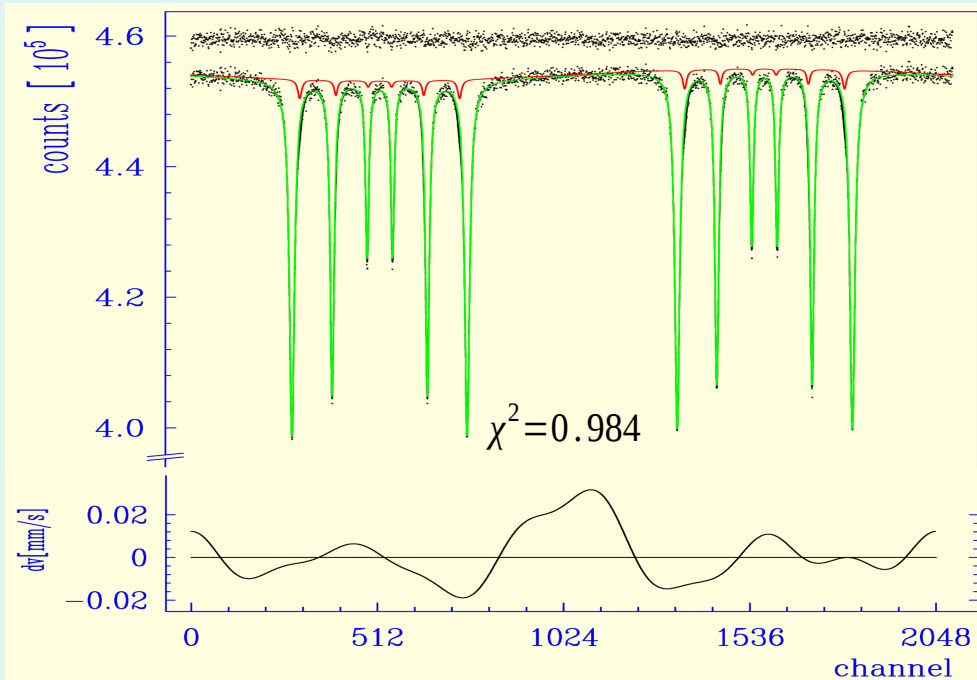


$$b_{fr} = \frac{A_b(\text{Ni})/0.9695}{A_t/0.9877} = 1.019 \frac{A_b(\text{Ni})}{A_t}$$

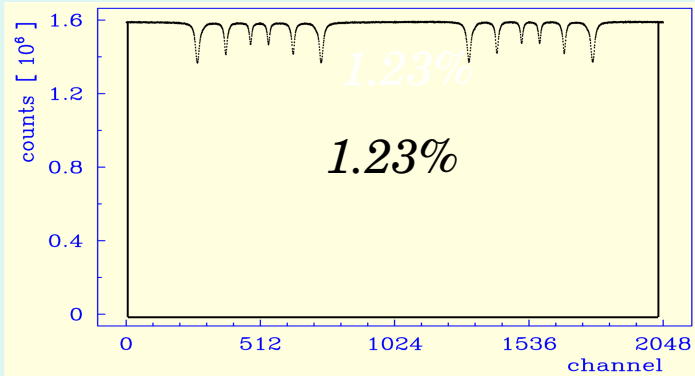
$\alpha$ -iron 25.5  $\mu\text{m}$ ,  $f=0.80$ ,  $\Gamma=\Gamma_N$

**Green:** 33.05 Tesla

**Red:** 30.67 Tesla (Mn)

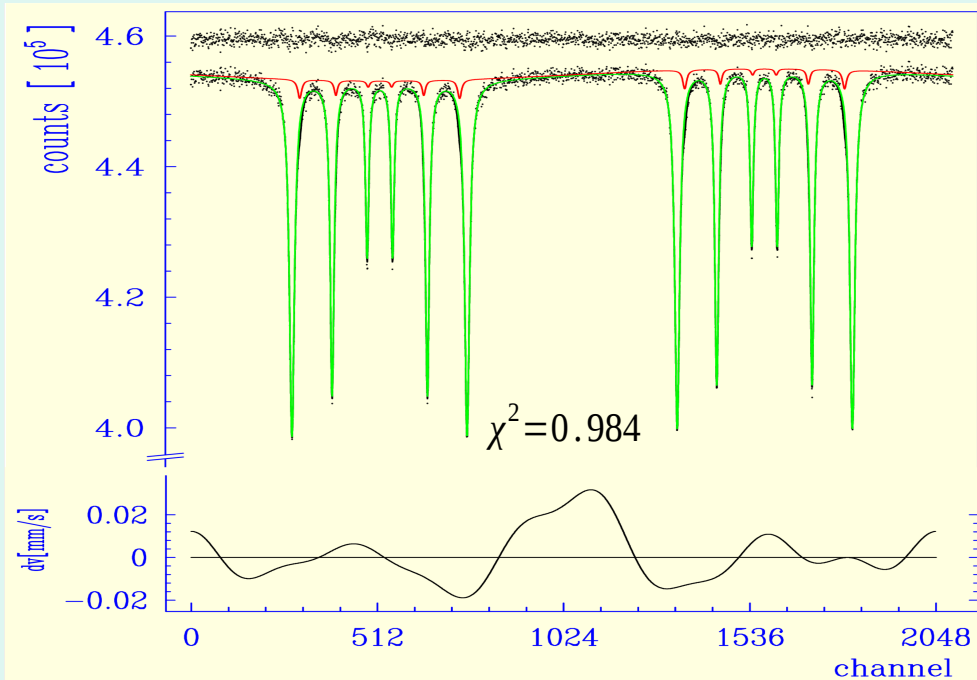


# Background



$$b_{fr} = \frac{A_b(\text{Ni})/0.9695}{A_t/0.9877} = 1.019 \frac{A_b(\text{Ni})}{A_t}$$

$\alpha$ -iron 25.5  $\mu\text{m}$ ,  $f=0.80$ ,  $\Gamma=\Gamma_N$



**Green: 33.05 Tesla**

**Red: 30.67 Tesla (Mn)**

**Fit results:**

$$f_{\text{source}}(1-b_{g_{fr}}) \Rightarrow$$

$$f_{\text{source}} = 0.715$$

$$\Gamma_{\text{source}} = 1.38 \cdot \Gamma_N$$

$$\sigma_{\text{source}} = 0.59 \cdot \Gamma_N$$

# ***Conclusion: some good practices***

## ***Taking down to the logbook:***

*Source (link to data sheet), Absorber (mg/cm<sup>2</sup>), Geometry (L<sub>0</sub>, aperture), drive(mode, frequency), counting system (detector, electronic settings)  
Background fraction (rate measurements: gated/ungated, Ni-foil), etc*

## ***Use of the convolution integral***

## ***Use of a reliable nonlinearity correction***

### ***Advantage:***

- Reliable values of  $t_{\text{eff}}$  instead relative areas  
(independent of the complexity of the applied theory)*
- Continuous control of the components of the equipment*